


Ab initio calculations to support accurate modelling of the rovibronic spectroscopy calculations of vanadium monoxide (VO)

Laura K. McKemmish, Sergei N. Yurchenko & Jonathan Tennyson

To cite this article: Laura K. McKemmish, Sergei N. Yurchenko & Jonathan Tennyson (2016) Ab initio calculations to support accurate modelling of the rovibronic spectroscopy calculations of vanadium monoxide (VO), *Molecular Physics*, 114:21, 3232-3248, DOI: [10.1080/00268976.2016.1225994](https://doi.org/10.1080/00268976.2016.1225994)

To link to this article: <http://dx.doi.org/10.1080/00268976.2016.1225994>

 View supplementary material 

 Published online: 15 Sep 2016.

 Submit your article to this journal 


 Article views: 85

 View related articles 

 View Crossmark data 

RESEARCH ARTICLE

Ab initio calculations to support accurate modelling of the rovibronic spectroscopy calculations of vanadium monoxide (VO)

Laura K. McKemmish , Sergei N. Yurchenko and Jonathan Tennyson

Department of Physics and Astronomy, University College London, London, UK

ABSTRACT

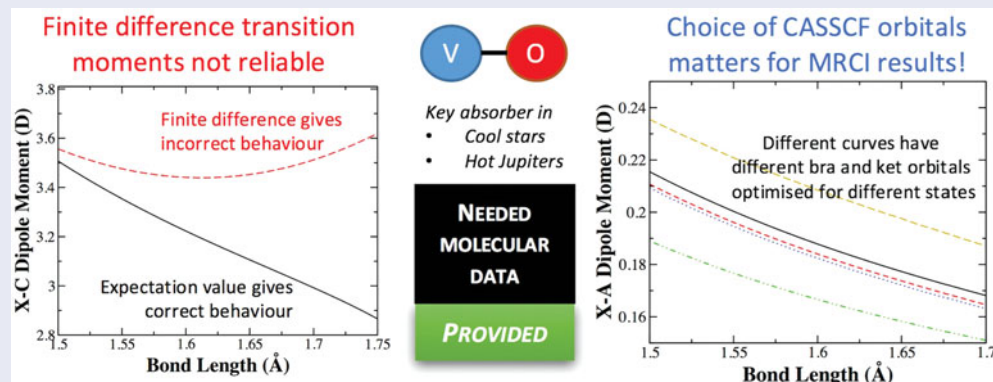
Accurate knowledge of the rovibronic near-infrared and visible spectra of vanadium monoxide (VO) is very important for studies of cool stellar and hot planetary atmospheres. Here, the required *ab initio* dipole moment and spin-orbit coupling curves for VO are produced. This data forms the basis of a new VO line list considering 13 different electronic states and containing over 277 million transitions. Open shell transition, metal diatomics are challenging species to model through *ab initio* quantum mechanics due to the large number of low-lying electronic states, significant spin-orbit coupling and strong static and dynamic electron correlation. Multi-reference configuration interaction methodologies using orbitals from a complete active space self-consistent-field (CASSCF) calculation are the standard technique for these systems. We use different state-specific or minimal-state CASSCF orbitals for each electronic state to maximise the calculation accuracy. The off-diagonal dipole moment controls the intensity of electronic transitions. We test finite-field off-diagonal dipole moments, but found that (1) the accuracy of the excitation energies were not sufficient to allow accurate dipole moments to be evaluated and (2) computer time requirements for perpendicular transitions were prohibitive. The best off-diagonal dipole moments are calculated using wavefunctions with different CASSCF orbitals.

ARTICLE HISTORY

Received 6 July 2016
Accepted 8 August 2016

KEYWORDS

Ab initio; spectroscopy; MRCI; transition metal diatomic; VO





1. Introduction

Transition metal diatomics are important absorbing species in many high temperature systems, particularly cool stellar [1] and hot planetary atmospheres [2], and in industrial processes, such as biomass gasification [3] and the incineration stage of waste disposal [4], where contamination by heavy metals, particularly mercury, is of significant environmental concern. Here, we focus on the

vanadium monoxide (VO) molecule, which is of particular interest to astronomers modelling late M-type stars and hot Jupiter exoplanets. The potential energy curves of the low-lying electronic states of VO are shown in Figure 1 for reference. Here, we report on our methodology explorations, and detail the methodology used to produce the final *ab initio* data.

A significant number of quantum chemistry calculations of VO have been performed previously [5–13].

CONTACT Laura K. McKemmish  l.mckemmish@ucl.ac.uk

 Supplemental data for this article can be accessed at:  <http://dx.doi.org/10.1080/00268976.2016.1225994>.

© 2016 Informa UK Limited, trading as Taylor & Francis Group

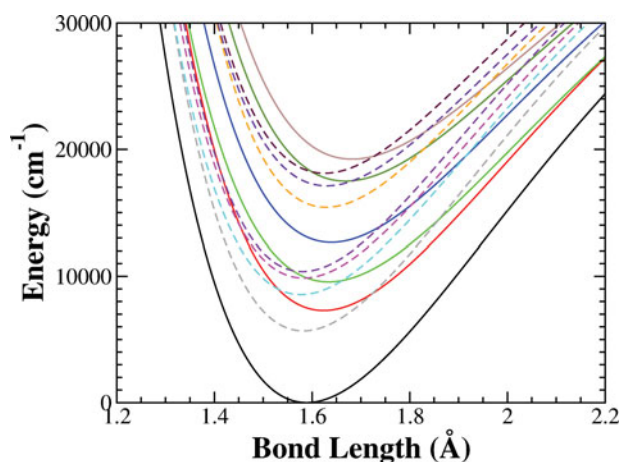


Figure 1. VO potential energy curves considered in this work. Curves in ascending order are: solid; $X^4\Sigma^-$, $A'^4\Phi$, $A^4\Pi$, $B^4\Pi$, $C^4\Sigma^-$, $D^4\Delta$; dashed; $a^2\Sigma^-$, $b^2\Gamma$, $c^2\Delta$, $d^2\Sigma^+$, $e^2\Phi$, $f^2\Pi$, $g^2\Pi$.

For equilibrium properties and the lowest electronic state of a given spin-symmetry, single-reference methods can give good results, e.g. the coupled-cluster singles, doubles and perturbative triples, CCSD(T), calculations of Bauschlicher *et al.* [5]. However, for good surfaces (particularly longer bond lengths) and other electronic states, multi-reference methods are imperative. Fortunately, computational advances have allowed studies using the higher quality internally-contracted multi-reference configuration interaction (icMRCI) methodology [6,14,15]. Milordos *et al.* [14] performed a thorough *ab initio* study of the equilibrium properties of the nine lowest electronic states of VO using MRCI methods. Milordos *et al.* used a state averaged approach to obtain results for the excited states; however, it is unclear which states were included in each calculation. Total energies, equilibrium bond lengths, electronic excitation energies, vibrational energies and equilibrium dipole moments were given for all states, including comparison to experimental data where available. Equilibrium values of the spin-orbit coupling constants are found for the triplet non- Σ states. A 2015 study by Hübner *et al.* [15] calculates the energetics of a much larger number of electronic states. However, no dipole moment or spin-orbit couplings were reported, except for the ground-state equilibrium dipole moment. Hübner *et al.* investigated the effect of including 3p correlation on their icMRCI results. Their calculations were much more computationally demanding but did not produce significant improvement in the potential energy curves. In particular, Hübner *et al.*'s icMRCI calculations fail to predict the correct ordering of the $C^4\Sigma^-$ and $D^4\Delta$ states.

Other 3d transition-metal/first-row-atom diatomic species have proven to have similar challenges [16–34]. These studies illustrate the power of the MRCI approach, which has consistently been shown to be a reasonably robust way of investigating low-lying excited states of these systems to usually at least semi-quantitative accuracy when compared against experiment. However, the studies also demonstrate the limitations of even these very high accuracy methods, e.g. correct modelling of higher electronic states and quantitative results. Particularly significant is the large errors found in excitation energies, often in excess of 1000 cm^{-1} [35]. These issues have stimulated new methodological developments in multi-reference methods, e.g. density-matrix renormalisation group theory [36–40], Monte Carlo configuration interaction [41], and stochastic multi-reference self-consistent field [42].

We produce the *ab initio* data required for a detailed model of the rovibronic spectroscopy of this molecule. The most important *ab initio* quantity in the spectroscopic model is the off-diagonal dipole moment, as there is no experimental reference. Note that potential energy curves and spin-orbit coupling curves, though important, can be empirically corrected based on experimental VO results. The diagonal dipole moment curves for VO will produce the rotational and vibrational transition intensities but these are not as critical for astronomical purposes as the electronic spectra.

In this paper, our methodological discussion focuses on the effect of:

- choice of states used to optimise the complete active space self-consistent field (CASSCF) orbitals on the icMRCI calculation energies, dipole moment, electronic angular momentum and spin-orbit coupling matrix elements,
- finite-field difference vs. expectation value off-diagonal dipole moments,
- using different vs. identical CASSCF orbitals for off-diagonal dipole moments and electronic angular momentum matrix elements and
- inclusion or exclusion of Davidson correction on the final properties

We use icMRCI calculations based on CASSCF orbitals; however, many of our conclusions can be expected to apply, or be extrapolated easily, to other similar multi-reference correlated methods, including the very popular complete active space perturbation theory level 2 (CASPT2) [43]. In this paper, we draw conclusions on best practice in making the above choices. This will be

based on theoretical principles, supported by experimental data.

Further, it is extremely useful to be able to gauge the accuracy of a particular theoretical calculation, i.e. to quantify its uncertainty, even in a non-rigorous approximate way [44]. We suggest that a good way of doing this in to evaluate the sensitivity of the given quantity to changed methodology (where the different methodology is slightly inferior on theoretical grounds, but not considerably so). This uncertainty quantification will be useful information to those who use our data, such as experimentalists and astronomers.

2. Quantitative study of *ab initio* methodology

2.1. Method

In this section, we will investigate the properties of VO using a few different methodologies in order to assess the best methodology choices, as well as estimate the uncertainty of the results. Our methodologies are chosen to be reasonably modest in cost because we are obtaining curves rather than single points, and we require results for 13 electronic states.

In view of time constraints, previous investigations and to focus on current studies, we keep some methodology constant. Unless otherwise specified, all calculations use aug-cc-pVQZ basis set. We use the CASSCF [45] implementation within MOLPRO [46] to find the molecular orbitals that are then used in MOLPRO's ic-MRCI [47] program. Note we are using the newly recommended notation [35] icMRCI(*n*) to indicate *n* states have been requested in the calculation.

We use a full-valence (3d, 4s/V, 2s, 2p/O) active space. Core-correlation effects are quite expensive to include and have been investigated elsewhere [14]. Therefore, we do not consider them in this manuscript. As discussed by Tennyson [48], the effects of core correlation and relativistic effects often partially cancel each other in practice. Thus, we will not consider relativistic corrections in this manuscript in the sense that we do not calculate Darwin and mass-velocity terms or use the Douglas–Kroll Hamiltonian etc. However, we do calculate spin–orbit coupling.

State-averaged CASSCF is often used for *ab initio* electronic spectroscopy, commonly without reference to the states used in the state-averaging. However, the large number of electronic states considered for our VO spectroscopic model led us to consider the influence of the way in which the CASSCF orbitals are optimised on the final icMRCI answer. Theoretically, state-specific calculations should give superior results as they are optimised for each individual electronic state (assuming that dynamic

correlation introduced post-CASSCF is small). However, the magnitude of this effect is unclear for VO *a priori* and is investigated in this section.

The standard method for evaluating dipole moments is via an expectation value (XP) of the wavefunction over the dipole moment operator. However, it has been found that finite-field difference (FD) expressions can be more accurate for diagonal dipole moments. The FD expression for a diagonal dipole moment is given by

$$\langle \Psi | \hat{\mu} | \Psi \rangle = \frac{E(+F) - E(-F)}{2F} \quad (1)$$

where *F* is the strength of the finite electric field, and *E*(*F*) is the energy of the molecule in the presence of the electric field. FD dipoles generally have better convergence properties than XP dipoles [49], and are usually closer to experimental values, see Rendell *et al.* [50] and references therein.

Almost exclusively, the expectation value expression is used to evaluate off-diagonal dipole moments. However, the finite-field different methodology is also an option. The mathematics were established by Adamson *et al.* [51], who also discusses a small set of test cases. Specifically, in symmetried form, we use

$$\langle \Psi_1 | \hat{\mu} | \Psi_2 \rangle = \frac{E(\Psi_1) - E(\Psi_2)}{4F} \times \frac{\langle \Psi_1(+F) | \Psi_2(-F) \rangle - \langle \Psi_1(-F) | \Psi_2(+F) \rangle}{4F} \quad (2)$$

where $\Psi(F)$ indicates the wavefunction in the present of the electric field *F*. This FD method for off-diagonal dipole moments is more complicated than for diagonal dipole moments and raise other issues which are discussed below.

2.2. Results

2.2.1. Potential energy curves

Several parameters characteristic of the quartet potential energy curves (PECs) (term energies, *T_e*, experimental vibrational frequencies, *T₁*, rotational frequencies, *r_e* etc.) are known from experiment [52–59]. This makes the PECs a useful tool for assessing different levels of theory. The excitation energies and vibrational frequencies are taken from the model Hamiltonian fits for each vibronic band in the original experimental papers.

2.2.1.1. Absolute minimum energies. Quantitative results are given in Table 1. Energies (usually *E* or *E_{min}*) are given relative to the energy of the X state using icMRCI(1)/aug-cc-pVQZ with no relativistic corrections, no core correlation and using orbitals from (X)-CAS. If

Table 1. Quantitative comparison of the quartet state results using different orbitals with icMRCI(n)/aug-cc-pVQZ calculations using (S)-CAS orbitals; $n = 1$ for X, A', and D states, 2 for the A state and 3 for the B and C states. Equilibrium bond lengths for the X, A', A, B, C and D states are 1.59, 1.63, 1.64, 1.67 and 1.69 Å, respectively. Energies and spin-orbit couplings are in cm^{-1} , dipole moments in units of D and the derivatives are taken in units of Å^{-1} . Experimental values are r_0 and T_0 .

S	\bar{E}_{min}	$\bar{E}_{\text{min}}^{+Q}$	r_e	T_1	T_2	r_e^{+Q}	T_1^{+Q}	T_2^{+Q}	μ_e^{XP}	μ_e^{FD}	$\mu_e^{\text{FD}(+Q)}$	μ_e^{XP}	μ_e^{FD}	$\mu_e^{\text{FD}(+Q)}$	SO'_e	SO_e	SO'_e	
$X^4\Sigma^-$	0	0	1.583	1028.5	2046.5	1.586	1014.7	2018.6	-2.50	-2.94	-3.08	-6.13	-6.54	-6.66				
	918	174	1.587	1025.4	2040.7	1.587	1021.1	2031.9	-2.41	-2.81	-2.98	-5.21	-6.09	-6.29				
	1053	160	1.588	1017.3	2024.6	1.587	1017.6	2025.3	-2.58	-2.88	-3.05	-5.07	-6.19	-6.63				
	1476	319	1.587	1031.2	2052.9	1.587	1019.6	2029.9	-2.39	-2.90	-3.03	-3.03	-6.12	-6.97				
	1382	550	1.588	1021.5	2032.6	1.588	1013.6	2017.1	-2.01	-2.69	-2.84	-6.14	-5.92	-6.33				
	1146	191	1.588	1016.7	2023.5	1.587	1019.2	2028.8	-2.59	-2.87	-3.04	-4.94	-6.13	-6.61				
	1555	549	1.588	989.2	2006.0	1.587	997.8	1983.0	-2.26	-2.78	-2.99	-5.45	-8.69	-6.69				
			1.599	983*		1.605	980*		-2.61	-3.09	-3.22							
			1.589	997*		1.592	992*		-2.43	-3.00	-3.23							
			1.588	1006*		1.591	1002*		-2.52	-3.10	-3.26							
			1.597	1006#														
			1.591	1020#														
			1.592	1001.8														
		0	0	1.5920	1001.8		1.589	1001.8		-3.36	-3.36	-3.36						
	$A^4\Phi$	9854	8432	1.633	941.4	1871.0	1.629	951.0	1890.7	-2.78	-3.28	-3.37	-5.06	-5.51	-5.95	256.3	-47.2	
9873		8443	1.633	941.8	1872.1	1.629	951.6	1892.2	-2.77	-3.28	-3.37	-5.05	-5.48	-5.91	254.1	-45.9		
10,045		8515	1.632	935.2	1858.3	1.628	948.2	1885.6	-2.90	-3.34	-3.43	-4.90	-5.58	-6.24	237.9	-37.0		
10,610		8909	1.636	933.3	1857.2	1.630	951.5	1894.1	-2.59	-3.28	-3.38	-5.10	-5.61	-6.20	240.4	-28.6		
10,200		8539	1.636	932.9	1843.7	1.630	946.7	1875.8	-2.75	-3.32	-3.40	-5.14	-5.43	-5.89	252.1	-38.4		
10,039		8497	1.634	937.7	1862.7	1.630	949.7	1887.3	-2.75	-3.31	-3.39	-5.11	-5.44	-5.89	241.8	-30.6		
10,076		8538	1.633	934.6	1857.5	1.629	947.4	1884.3	-2.86	-3.35	-3.44	-4.95	-5.60	-6.21	239.7	-33.8		
9821		8207	1.631	942.7	1875.0	1.627	955.5	1901.8	-2.86	-3.34	-3.43	-4.82	-5.68	-6.28	240.0	-37.0		
10,339		8572	1.633	853.6	1747.4	1.627	892.0	1817.1	-2.61	-3.31	-3.40	-4.89	-6.07	-6.43	241.8	-30.6		
7187		7084	1.639	913*		1.640	912*		-3.25	-3.59	-3.60							
8049		7753	1.628			1.638			-3.01									
7956		7635	1.627			1.639												
7598%			1.635	944#														
7816%			1.630	955#														
7565%			1.634	950#														
	7255	7255	1.629	936.5	1865.0	1.629	936.5	1865.0										
$A^4\Pi$	12,530	10,999	1.641	905.5	1797.9	1.636	920.3	1828.5	-2.73	-3.26	-3.35	-4.82	-5.37	-5.89	69.0	55.1		
	12,519	10,988	1.641	904.8	1796.6	1.635	919.7	1827.1	-2.74	-3.26	-3.35	-4.82	-5.38	-5.90	69.9	55.2		
	12,600	11,014	1.641	898.3	1783.5	1.636	915.1	1818.9	-2.91	-3.40	-3.50	-4.77	-5.86	-6.68	55.5	72.4		
	13,017	11,224	1.646	902.3	1796.4	1.639	922.0	1837.8	-2.60	-3.32	-3.41	-4.92	-5.84	-6.51	60.7	66.8		
	12,665	11,030	1.643	900.4	1786.8	1.636	917.2	1821.7	-2.72	-3.28	-3.37	-4.88	-5.34	-5.88	61.4	69.0		

(Continued)



Table 1. Continued

S	\tilde{E}_{\min}	$\tilde{E}_{\min}^{\text{H-Q}}$	r_e	T_1	T_2	$r_e^{\text{H-Q}}$	$T_1^{\text{H-Q}}$	$T_2^{\text{H-Q}}$	μ_e^{XP}	μ_e^{FD}	$\mu_e^{\text{FD}(\pm)}$	μ_e^{XP}	μ_e^{FD}	$\mu_e^{\text{FD}(\pm)}$	SO_e	SO_e	SO_e
(X,A',A,B)	12,636	11,042	1.642	898.0	1783.2	1.636	914.4	1817.7	-2.86	-3.39	-3.49	-4.87	-5.89	-6.64	58.1	68.2	68.2
(X,A',A,B,C)	12,454	10,860	1.642	902.2	1796.7	1.638	913.6	1821.6	-2.86	-3.38	-3.48	-4.64	-5.85	-6.58	57.5	72.9	72.9
(X,A',A,B,C,D)	12,926	11,195	1.637	1008.9	1944.2	1.636	960.5	1880.1	-2.62	-3.34	-3.42	-4.68	-6.32	-6.72	61.4	69.0	69.0
MRCI/B ^a	9821	9614	1.650	866*		1.651	866*		-3.62	-3.84	-3.90						
C-MRCI/BP ^a	10,676	9677	1.638			1.639			-3.57								
C-MRCI+DKH2/BP ^a	10,669	9644	1.637			1.639			-3.55								
(A,A,B)-CAS, icMRCI ^b	10138%		1.644	906 [#]													
(A,A,B)-CAS, C-MRCI ^b	10,259%		1.639	920 [#]													
(A,A,B,3 ⁴ T)-CAS, icMRCI ^b	10,106%		1.644	910 [#]													
Exp[53]	9499	9499	1.637	884.0		1.638	884.0										
B ⁴ Π																	
(A',A,B)	17,241	13,151	1.658	937.1	1879.3	1.643	952.2	1907.7	-6.39	-6.64	-5.99	-3.22	-7.85	-8.74	90.6	-35.6	-35.6
(A',A,B,D)	18,622	13,938	1.652	1051.8	2127.7	1.634	1054.7	2132.1	-6.24	-7.35	-6.85	-0.84	-6.01	-7.00	89.1	-30.2	-30.2
(X,A',A,B)	17,630	13,230	1.659	951.8	1914.0	1.641	968.7	1943.6	-6.30	-6.85	-6.20	-3.34	-9.24	-10.73	89.8	-30.4	-30.4
(X,A',A,B,C)	17,635	13,251	1.656	973.1	1966.0	1.641	984.1	1982.7	-6.40	-6.88	-6.23	-3.36	-9.74	-11.40	90.8	-29.2	-29.2
(X,A',A,B,C,D)	18,790	13,967	1.650	1153.8	2276.9	1.633	1126.0	2212.8	-6.25	-7.53	-7.05	-1.85	-17.42	-21.17	89.6	-32.0	-32.0
MRCI/B ^a	14,743	12,864	1.662	870*		1.654	894*		-6.45	-6.39	-5.74						
C-MRCI/BP ^a	14,889	13,491	1.652			1.644			-6.75								
C-MRCI+DKH2/BP ^a	15,839	14,460	1.650	1006*		1.643	1002*		-6.82								
(A,A,B)-CAS, icMRCI ^b	14,268%		1.658	897 [#]													
(A,A,B)-CAS, C-MRCI ^b	15,792%		1.658	904 [#]													
(A,A,B,3 ⁴ T)-CAS, icMRCI ^b	14,115%		1.654	913 [#]													
Exp[56,57]	12,606	12,606	1.644	901.0		1.644	901										
C ⁴ Σ ⁻																	
(X,C)	19,918	17,222	1.665	973.4	1974.9	1.665	959.9	1947.8	-4.47	-4.70	-4.52	-5.82	-13.40	-13.84			
(X,C,D)	20,985	17,635	1.677	887.6	1769.8	1.667	895.2	1790.7	-4.50	-5.27	-4.74	-4.00	-10.36	-12.11			
(X,C,D)-CAS, icMRCI ^b	17,712%		1.683	851 [#]													
(X,C,D)-CAS, C-MRCI ^b	19,285%		1.675	872 [#]													
(X,C,D,3 ⁴ Σ ⁻ , 4 ⁴ Σ ⁻ , 1 ⁴ Π)-CAS, icMRCI ^b	17,091%		1.682	854 [#]													
Exp[52]	17,420	17,420	1.675	852.8	1699.6	1.675	852.8	1699.6									
D ⁴ Δ																	
(D)	18,300	17,310	1.679	834.2	1645.6	1.679	835.7	1649.6	-1.31	-1.25	-1.14	-5.48	-6.41	-6.57			
(X,C,D)	18,554	17,240	1.677	823.3	1603.5	1.677	833.4	1633.2	-1.17	-1.24	-1.14	-5.69	-5.85	-5.96			
(A',A,B,D)	20,564	18,033	1.699	782.6	1513.9	1.674	855.2	1689.7	-1.61								
(X,C,D)-CAS, icMRCI ^b	15,849%		1.683	850 [#]													
(X,C,D)-CAS, C-MRCI ^b	15,986%		1.672	863 [#]													
(X,C,D,3 ⁴ Σ ⁻ , 4 ⁴ Σ ⁻ , 1 ⁴ Π)-CAS, icMRCI ^b	15,462%		1.683	855 [#]													
Exp[54]	19,148	19,148	1.686	836.0		1.686	836										

^a *Ab initio* results from Miliordos and Mavridis [14].

^b *Ab initio* results from Hübner et al. [15].

* $\Delta G_{1/2}$

$\tilde{\nu}_e$

% T_e

rather than absolute energy above (X)-CAS, icMRCI(1)/aug-cc-pVQZ minimum.

Table 2. Off-diagonal dipole moment matrix elements evaluated using icMRCI(n)/aug-cc-pVQZ using (Bra S)-CAS for bra wavefunction and (Ket S)-CAS orbitals for the ket wavefunction; n is 1 for X, 2 for A and C, 3 for B. The equilibrium bond length is taken as 1.59 Å in all cases, for ease of comparison. Dipole moments in D and derivatives taken in units of D/Å⁻¹.

	Bra S	Ket S	μ_e	μ_e'
X-A	(X)	(A)	0.48	-0.60
X-A	(X)	(A',A)	0.47	-0.59
X-A	(X)	(A',A,B)	0.65	-0.04
X-A	(X)	(X,A)	0.44	-0.50
X-A	(X)	(X,A',A)	0.46	-0.55
X-A	(X)	(X,A',A,B)	0.54	-0.61
X-A	(X,A)	(A)	0.47	-0.58
X-A	(X,A)	(A',A)	0.46	-0.57
X-A	(X,A)	(A',A,B)	0.53	-0.59
X-A	(X,A)	(X,A)	0.43	-0.48
X-A	(X,A)	(X,A',A)	0.45	-0.53
X-A	(X,A)	(X,A',A,B)	0.52	-0.59
X-A	(X,A',A)	(A)	0.47	-0.57
X-A	(X,A',A)	(A',A)	0.45	-0.56
X-A	(X,A',A)	(A',A,B)	0.53	-0.58
X-A	(X,A',A)	(X,A)	0.42	-0.47
X-A	(X,A',A)	(X,A',A)	0.44	-0.52
X-A	(X,A',A)	(X,A',A,B)	0.51	-0.58
X-A	(X,A',A)	(X,A',A)	0.49	-0.54
X-A	(X,A',A,B)	(A',A,B)	0.65	-0.05
X-A	(X,A',A,B)	(X,A',A,B)	0.70	0.00
X-A	(X,A',A,B,C)	(X,A',A,B,C)	0.64	-0.16
X-A	(X,A',A,B,C,D)	(X,A',A,B,C,D)	0.73	-0.21
X-B	(X)	(A',A,B)	1.51	-2.86
X-B	(X)	(X,A',A,B)	1.52	-2.82
X-B	(X,A)	(A',A,B)	1.53	-2.83
X-B	(X,A)	(X,A',A,B)	1.53	-2.78
X-B	(X,A',A,B)	(A',A,B)	1.51	-2.83
X-B	(X,A',A,B)	(X,A',A,B)	1.67	-3.12
X-B	(X,A',A,B,C)	(X,A',A,B,C)	1.55	-2.86
X-B	(X,A',A,B,C,D)	(X,A',A,B,C,D)	1.67	-3.55
X-C	(X)	(X,C)	2.92	-3.76
X-C	(X)	(X,C,D)	2.88	-2.93
X-C	(X,C)	(X,C)	3.03	-3.11
X-C	(X,C)	(X,C,D)	2.99	-2.20
X-C	(X,C,D)	(X,C,D)	3.08	-2.70
A-B	(A)	(A',A,B)	0.17	-1.23
A-B	(A)	(A',A,B,D)	0.10	-1.87
A-B	(X,A)	(A',A,B)	0.19	-1.24
A-B	(X,A)	(X,A',A,B)	0.22	-1.01
A-B	(A',A,B)	(A',A,B)	0.38	-0.59
A-B	(A',A,B,D)	(A',A,B,D)	0.39	-0.64
A-B	(X,A',A,B)	(X,A',A,B)	0.38	-0.48
A-B	(X,A',A,B,C)	(X,A',A,B,C)	0.37	-0.45
A-B	(X,A',A,B,C,D)	(X,A',A,B,C,D)	0.38	-0.49

the Davidson correction is used, then the energy is given relative to the X state energy from icMRCI(1)+Q/aug-cc-pVQZ with (X)-CAS.

The icMRCI calculation using CASSCF orbitals optimised for fewer states generally gives the lowest icMRCI minimum energy without the Davidson correction. There are some subtleties, e.g. the X state energy for (X,C)-CASSCF is higher than using (X,C,D)-CASSCF; however, this is compensated by the rise in C state energy.

With the Davidson correction, the general principle that fewer states in the CASSCF orbitals means lower icMRCI energies still largely holds true, though the

results are more mixed and, of course, the calculations no longer strictly obey the variational principle. For the X and A' state, the Davidson correction also dramatically decreases the difference in minimum energy arising from different CASSCF orbitals; without the Davidson correction, the difference in minimum energy is up to 1476 cm⁻¹, while with the Davidson correction, this decreases to 549 cm⁻¹. The Davidson correction increases the magnitude of the correlation included. This indicates that post-CASSCF correlation energy is larger when the CASSCF orbitals are less optimised for the particular state.

2.2.1.2. Excitation energies. By far the biggest error from pure *ab initio* calculations is the electronic excitation energies, i.e. T_0 . Without the Davidson correction, results in error of more than 5000 cm⁻¹ are observed; the *ab initio* T_0 is always larger than the experimental T_0 . The use of the Davidson correction significantly helps reducing errors to generally 1000–2000 cm⁻¹ though T_0 is still almost always overestimated. Clearly, this is far from spectroscopic accuracy.

It is very important to note that with different CASSCF orbitals for the ground and excited, T_0 is no longer uniquely defined. However, they can be inferred by taking differences between the minimum energies. Different choices of CASSCF orbitals do affect the result quite significantly; however, this does not account for our large errors.

Thus, there is something fundamentally missing from the current *ab initio* treatment of VO that causes the higher excited states to be less well described than the ground state (resulting in T_0 being overestimated). The effect of the Davidson correction indicates that inadequate treatment of electron correlation may be partially responsible.

2.2.1.3. Vibrational frequencies. The frequency of the transitions between the $v = 0$ and $v = 1$, and $v = 0$ and $v = 2$ vibrational levels are given as T_1 and T_2 , respectively. Note that we compare directly with the observed vibrational frequencies, not the harmonic frequencies; this is done by performing a single-state nuclear motion calculation for the vibrational energy levels using Duo [60].

The fundamental and first overtone vibrational frequencies vary by up to 20 cm⁻¹ for the X, A', A and B states using different CASSCF orbitals. Much larger errors, however, are seen for the B, C and D states, up to 100 cm⁻¹.

The T_1 and T_2 change when the Davidson correction is included, by up to 30 cm⁻¹. This change in

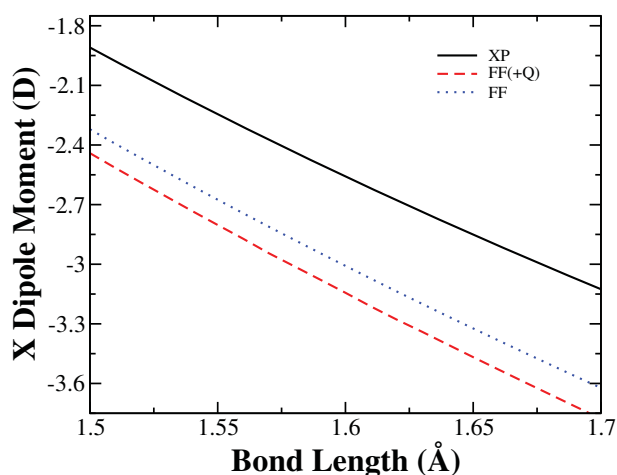


Figure 2. VO $X^4\Sigma^-$ state diagonal dipole moments, computed using icMRCI(1)/aug-cc-pVQZ with (X)-CAS.

vibrational energy is fairly systematic within a particular electronic state across the different choice of CASSCF orbitals. The non-Davidson corrected frequencies seem to match experiment better. However, both are generally higher than the experimentally observed frequencies.

2.2.1.4. Rotational frequencies. Equilibrium bond distances are symbolised by r_e . When combined with the reduced mass of a particular isotopologue, these largely determine the rotational frequencies. Table 1 shows that r_e shows variations on the order of 3 mÅ with respect to changing CASSCF orbitals. For large J and rovibronic transitions between electronic states with different errors in the rotational constants, this can have a significant influence on the final spectra. However, this is normally not an issue as the rotational constants are usually well defined from experiment.

2.2.2. Diagonal dipole moments

Figure 2 shows that for the $X^4\Sigma^-$ state, the FD diagonal dipole moment using (X)-CAS orbitals, both with the Davidson correction (-3.08 D at equilibrium) and without (-2.94 D at equilibrium), is significantly more negative than the XP one (-2.50 D at equilibrium). The experimental value is -3.355 ± 0.014 D [55]; the FD results are closer to this value. Note that this comparison is not strictly fair, as the experimental dipole moment is a vibrationally averaged quantity which we compare to the static equilibrium dipole, μ_e ; however, this should be reliable to within the errors discussed here. Further, spin-orbit coupling means higher excited states (and thus their dipole moments) could contribute to the lowest rovibronic state. However, since the X state is well separated from all other electronic states, this effect is very small.

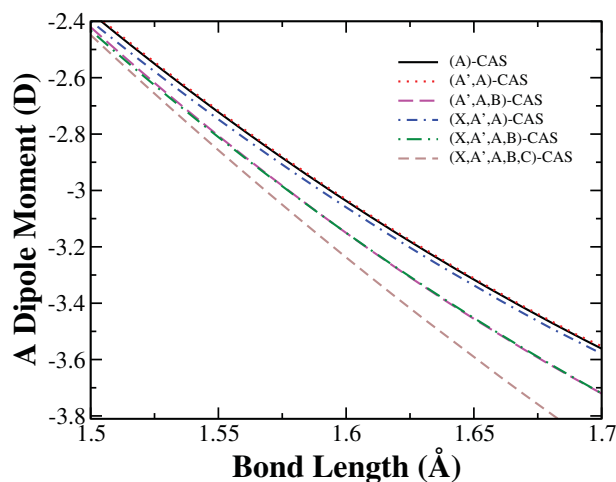
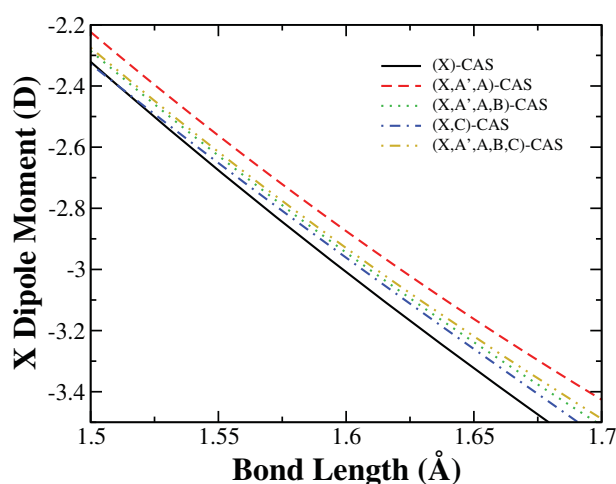


Figure 3. VO $X^4\Sigma^-$ (top) and $A^4\Pi$ (bottom) state diagonal dipole moment, evaluated with FD methodology, using icMRCI(1)/aug-cc-pVQZ and icMRCI(2)/aug-cc-pVQZ energies, respectively.

Figure 3 shows the difference in the DMC for the $X^4\Sigma^-$ state when evaluated using different CASSCF orbitals, with FD methodology and icMRCI(1)/aug-cc-pVQZ energies. The curves all have the same general shape, but they are off-set vertically and there is some deviation from parallel curves. Figure 3 also shows the difference in the DMC for the $A^4\Pi$ state when evaluated using different CASSCF orbitals, with FD methodology and icMRCI(2)/aug-cc-pVQZ energies. Apart from the CASSCF orbitals averaged over more than five states, the results separate clearly into two curves, based on whether the B state is included in the CASSCF orbital optimisation. From this result, it is apparent that the A and B state have significantly different charge distributions.

Table 1 compares the quartet states diagonal dipole moments calculated using different orbitals. Generally different orbitals give dipole moments that vary by up to 15%. The derivatives have larger variation, up to approximately 30%. The finite-field dipole moments and derivatives are generally about 0.5 D and 0.5 D/Å, respectively,

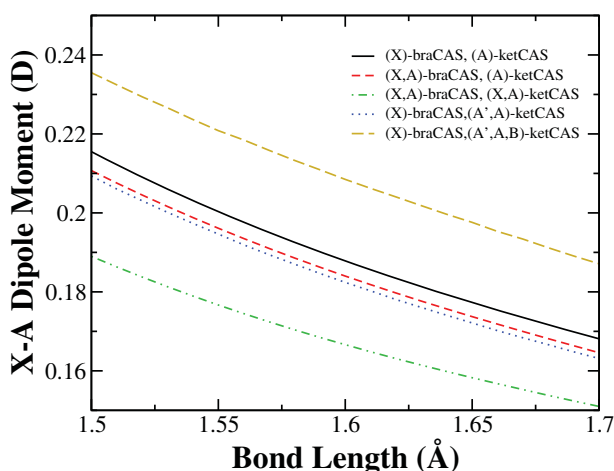


Figure 4. X-A off-diagonal dipole moment, bra wavefunction (of $X^4\Sigma^-$ state) from icMRCI(1)/aug-cc-pVQZ with braCAS orbitals, ket wavefunction (of $A^4\Pi$ state) from icMRCI(2)/aug-cc-pVQZ with ketCAS orbitals.

lower than the expectation value dipole moments and derivatives.

2.2.3. Off-diagonal dipole moments

2.2.3.1. Finite field. We report calculations of the off-diagonal FD dipole moments compared to XP for VO in Table 3, giving details of both numerical results and timings.

Parallel transitions

FD calculations for parallel transitions, typified by the X-C transition, are practical in terms of calculation time. The use of FD rather than XP triples the icMRCI calculation time.

However, it is unclear what is the best choice for the energies of the zero-field wavefunctions; with or without Davidson correction, with the same orbitals or different orbitals for each state, using experimental values for energies, etc. This has a substantial effect of the magnitude of the dipole moment as it is a pure multiplicative factor. Though one of these FD results may in fact be more accurate than the XP result, it is currently unclear which result this would be.

One further cause for concern for the off-diagonal FD dipole moments is consideration of their behaviour as a function of bond length, shown in Figure 5. The XP curve goes smoothly towards zero at long bond lengths (as expected for this transition). However, some FD results seem to increase with increasing bond length. The dipole moment may eventually turn over at long bond lengths, but this behaviour does not inspire confidence in the FD results.

Perpendicular transitions

For perpendicular transitions, the finite field needs to be applied perpendicular to the molecular axis which

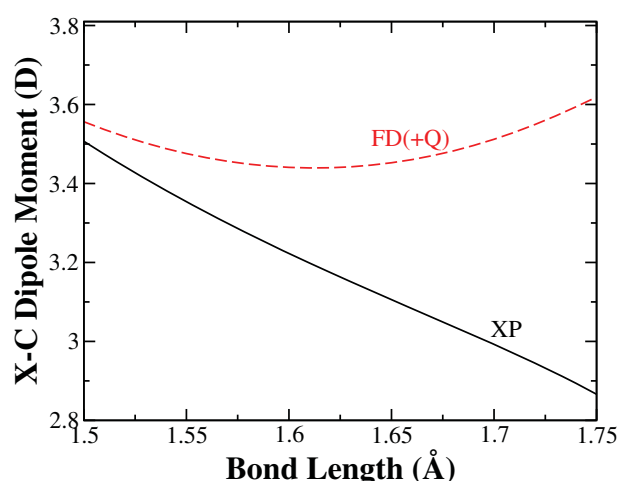


Figure 5. $X^4\Sigma^- - C^4\Sigma^-$ off-diagonal dipole moment, using icMRCI(2)/aug-cc-pVQZ with (X,C)-CAS for bra and ket wavefunction. The solid line is calculated using expectation value method, the dotted line is finite-field method using energies with Davidson correction.

breaks the linear-molecule symmetry. This means C_1 symmetry must be used: all electronic states of a given spin and both components of Π , Δ , Φ etc. states have to be included in the calculation. For example, for the X-A dipole, we need five states. Table 3 presents some CASSCF/aug-cc-pVDZ and icMRCI/cc-pVDZ results. For the X-A dipole moment, the reduction in symmetry increases the calculation time more than 25-fold for icMRCI(5) in C_1 vs. icMRCI(1,2,2,0) in C_{2v} symmetry for the XP method, and even more for the FD method (calculation times around 80 times longer), due to the difficulty of converging the larger number of states and increased number of reference states in the icMRCI.

2.2.3.2. Expectation value. Given the difficulties experienced with the finite-field off-diagonal dipole moments and the importance of this parameter to the final spectroscopic model, we investigate a range of different CAS orbitals here. Figure 4 shows the X-A off-diagonal dipole moment curves for five choices of bra and ket CAS orbitals calculated using the expectation value methodology. Table 4 provides quantitative comparison metrics for the X-A, X-B, X-C and A-B off-diagonal expectation value dipole moments.

In most cases, different orbital choices give qualitatively similar results for both the off-diagonal dipole moment and electronic angular momentum matrix elements. A variation of 10 % between the equilibrium off-diagonal dipole moments due to different CASSCF choices is usual. The derivatives are more variable with deviations of more than 40 % seen for different CASSCF choices in the X-C dipole moment derivative (-2.20 to -3.76 D/Å).

Table 3. Comparison of off-diagonal dipole moment of VO at a bond distance of 1.59 Å, using different methodologies as specified in text. Energies are given in cm⁻¹, dipole moments in D.

	Symmetry	Method	ΔE	μ_e	Time (min)	
X-A	$C_{\infty v}$	CASSCF(1,1,1,0)	XP	0.267	3.6	
	C_1	CASSCF(5)	XP	0.267	6.1	
	C_1	CASSCF(5)	FD	13,879.4	0.359	9.0
	C_1	CASSCF(5)+Exp.E	FD	9498.9	0.245	"
	C_{2v}	icMRCI(1,2,2,0)	XP		0.492	16.0
	C_1	icMRCI(5)	XP		0.507	421.9
	C_1	icMRCI(5)	FD	12,154.0	0.546	1321.4
	C_1	icMRCI(5)+Exp.E	FD	9499	0.427	"
	C_1	icMRCI(5)+Q	FD	10,931.55	0.491	"
	X-C	$C_{\infty v}$	CASSCF(2)	XP	2.779	2.8
$C_{\infty v}$		CASSCF(2)	FD	24,847.6	3.710	3.3
$C_{\infty v}$		CASSCF(2)+Exp.E	FD	17,420.1	2.6	"
C_{2v}		icMRCI(2)	XP		3.309	10.8
C_{2v}		icMRCI(2)	FD	19,397.1	3.712	27.0
C_{2v}		icMRCI(2)+Exp.E	FD	17,420.1	3.334	"
C_{2v}		icMRCI(2)+Q	FD	17,739.8	3.395	"

Table 4. Off-diagonal spin-orbit matrix element calculations using CASSCF/aug-cc-pVDZ using (S)-CAS. The equilibrium bond length is taken as 1.59 Å in all cases, for ease of comparison. Spin-orbit couplings are given in cm⁻¹ and derivatives taken in units of cm⁻¹/Å.

	S	SO _e	SO _e '
X-A	(X,A)	68.9	10.5
X-A	(X,A',A)	69.8	11.1
X-A	(X,A',A,B)	66	15.4
X-A	(X,A',A,B,C)	66.5	15.4
X-A	(X,A',A,B,C,D)	65.3	23.7
X-B	(X,A',A,B)	7.9	45.4
X-B	(X,A',A,B,C)	7.5	46.4
X-B	(X,A',A,B,C,D)	5.3	43.7
A-B	(A',A,B)	2.7	-2.8
A-B	(A',A,B,D)	3.2	-1.9
A-B	(X,A',A,B)	2.2	-2.7
A-B	(X,A',A,B,C)	2.2	-3
A-B	(X,A',A,B,C,D)	2.6	-3

However, there are some cases where there is a much larger, qualitatively significant, effect. In particular, it is clear from the magnitude of the X-A and A-B dipole moments and their derivatives that the inclusion of the B state in the CASSCF orbital optimisation has a large effect on the electron distribution in the final icMRCI A wavefunction. For example, the X-A dipole moment increases by more than 25 % and the A-B dipole moment doubles. The absolute dipole moment increases by approximately 0.2 D in each case. An even larger effect is seen in the derivatives of the X-A dipole moment which is reduced dramatically from about 0.5 D/Å to near zero when the B state is included. Again, we see that the B state must have a fundamentally different character to the A state.

2.2.4. Spin-orbit coupling parameters

Previous benchmarking on ScH [49] suggested that CASSCF/aug-cc-pVDZ calculations were sufficiently

accurate for the construction of line lists for TM diatomics. We follow this method and basis set, and focus on the influence of the electronic states included in the CASSCF calculation. Note that for spin-orbit coupling elements, the bra and ket wavefunctions must use the same CASSCF orbitals.

Table 1 compares the diagonal spin-orbit coupling matrix elements for all non-singlet quartet states. Table 4 provides comparison metrics of some off-diagonal quartet spin-orbit couplings matrix elements at equilibrium. Both these comparisons demonstrate that different choices of states in the CASSCF affects the matrix element by roughly 3–5 cm⁻¹ in absolute magnitude. This is similar to that observed for the variation between different basis sets and icMRCI vs. CASSCF for ScH [49].

2.3. Discussion

2.3.1. Davidson correction

The Davidson correction [61,62] (conventionally indicated by +Q) provides an estimate of the triples and quadruples contribution to the correlation energy, improving the size-consistency of the calculation. Previous icMRCI studies on related TM oxides show that the Davidson correction influences the excitation energy significantly, up to 2000 cm⁻¹ [14,23,29]; often (but not always) this is a decrease in energy that improves agreement with experiment. These studies show changes in bond length, harmonic vibrational frequency and diagonal dipole moment (computed using finite fields) are fairly small: about 0.005 Å, 10 cm⁻¹ and 0.1 D respectively. However, it is not clear whether the Davidson correction improves results comparing to experiment. Our results provide additional quantitative data on this effect, illustrating the generality of the previous results.

2.3.2. Choice of orbitals for multi-reference calculations

The results in this section demonstrate that the way in which the CASSCF orbitals are optimised, specifically what electronic states are considered, has a profound effect on the final icMRCI calculation. A particularly striking example seen in VO is that inclusion of the B state in optimising CASSCF orbitals for the icMRCI calculations of the X, A' and A states significantly affects the DMC.

On a semi-quantitative level, the property curves obtained using different orbitals provides an estimate of the uncertainty in the prediction of the property. The 'best' answer is taken as the result using the CASSCF orbitals optimised for the smallest number of electronic states.

From a theoretical perspective, generally optimising for as few states as possible is preferred as the CASSCF orbitals more closely match the wavefunction of the electronic state under consideration. This will usually give lower icMRCI energies and, presumably (but not certainly), better properties [63]. However, there are at least three cases where we believe that extra electronic states must be included.

- (1) For higher states of a particular spin and symmetry, the lower energy states should be included.
- (2) For icMRCI calculations and low-symmetry C_{2v} SA-CASSCF calculations, all states of lower energy with the same C_{2v} symmetry must be included.
- (3) Sometimes, the order of electronic states is not correctly reproduced by calculations; therefore, states of slightly higher energy can be useful or necessary.

Note that both components of the $\Lambda \neq 0$ states should be included in C_{2v} SA-CASSCF calculations, for example, A_1 and A_2 in case of the Δ symmetry.

Ideally, there would be no reason to include extra states. However, due to the high nonlinearity of the problem, CASSCF has substantial convergence problems [64] and discontinuities in the PEC and/or DMC are common [63,65–68]. Therefore, extra states in the CASSCF orbital optimisation also can be useful because

- (4) Increasing the SA-CASSCF space can provide better convergence [69]. For example, it can help avoid root flipping [70].
- (5) SA-CASSCF helps recovering the degeneracy at the dissociation limit [71].

Finally, there are some pragmatic reasons why SA-CASSCF may be preferred:

- (6) Running calculations based on a single SA-CASSCF orbitals for all properties of interest is faster in terms of human time set-up costs.

- (7) To link icMRCI calculations based on SS- or MS-CASSCF orbitals, it is useful (and sometimes essential) to run SA calculations to find the correct phases and signs of wavefunctions and matrix elements.

Even with SS- and MS-CASSCF orbitals, icMRCI calculations do not give satisfactory answers for many properties. Of particular concern is the very large errors in the electronic excitation energies, sometimes by up to 5000 cm^{-1} . Even with the Davidson correction, non-systematic over-estimates of about 1000 cm^{-1} are the norm. Furthermore, even the qualitative ordering of states is often wrong, e.g. the $C^4\Sigma^-$ and $D^4\Delta$ states in VO are often switched. Aside from the obvious effect that these incorrect excitation energies have on the band origins, there are also more subtle effects. In particular, perturbations between observed ('bright') and hidden ('dark') states are often seen as resonance interactions between electronic states via, for example, off-diagonal spin-orbit interactions. Characterising these resonances, for example, by assigning electronic and vibrational states to the dark state is often difficult experimentally. Unfortunately, it would appear that at present *ab initio* methods remain unable to help with this for molecules like VO. In contrast, for main-group molecules such as C_2 *ab-initio* methodologies are greatly facilitating the identification of new spectral bands, e.g. [72].

2.3.3. Dipole moments

Worse than the errors in excitation energy, which at least can be generally quantified, are the unknown errors in the dipole moment curves. For VO, only the ground experimental value is known. Lifetimes have been measured for certain states [73] which serve as a useful verification of *ab initio* results [74]; more of these measurements would be helpful.

There are various theoretical arguments [75–78] to expect that dipole moments computed using finite-field differences (FD) are superior to those obtained using expectation values (XP). However, Erzernor *et al.* [76] show that there are certain requirements that must be fulfilled for the FD to be superior; in particular, the optimal energy must be obtainable with only second-order corrections to the orbitals and CI coefficients etc. For diagonal dipole moments, these conditions are generally fulfilled and theoretical argument about the superiority of the FD dipoles agrees with what is found in practice in *ab initio* calculations. This is also true for the ground state of VO [14]. However, our results show that it is probably not so straightforward for off-diagonal dipole moments.

Our results for the diagonal dipole moments confirm the advantage of using the finite-field differences rather than expectation values; automating this process

within quantum chemistry packages with new keywords would facilitate the use of this methodology. Note that CFOUR [79] has such routines for coupled-cluster related methodologies.

Given the importance of the off-diagonal dipole moments to the quality of the final line list, it is important that we maximise the accuracy of the expectation value results. Finite-field methodology seems a logical choice; however, the success of the finite-field equation for diagonal dipole moments is not matched for off-diagonal dipole moments. There are two specific issues. First, should the same electronic structure method be used to calculate the energies and wavefunctions of the two states? Transition energies are notoriously badly predicted by theory for transition metal diatomics; errors of 1000 cm^{-1} or more are regularly encountered [35]. Without the Davidson correction (which improves the energy but does not correct the wavefunction), for VO we found errors in excess of 5000 cm^{-1} . Another option is to use the experimental excitation energy. Second, the application of a finite electric field for a perpendicular transition in a diatomic molecule (e.g. between a sigma and pi electronic state) reduces the symmetry of the calculation. This means that a larger number of states need to be included in the calculation. This significantly increases the calculation time and decreases the accuracy of the result.

Given these issues with finite-field off-diagonal dipole moments, we must use the expectation value method. However, we argue that a significant improvement can be obtained by representing the icMRCI wavefunction of both the bra and ket wavefunctions using different, optimised CASSCF orbitals for each electronic state. Without a higher accuracy 'near-exact' result, we cannot demonstrate that the result is always improved. However, we can and do demonstrate that the variation of the off-diagonal dipole moments obtained using different CASSCF orbitals routinely vary by approximately 10%, with more extreme errors of up to 100% found when the CASSCF orbitals are trying to represent two states with very different charge distributions (e.g. when considering the $A^4\Pi$ and $B^4\Pi$ state in VO).

3. Final calculations

3.1. Method

For the final calculation for the spectroscopic model of VO, the electronic structure calculations were generally icMRCI/aug-cc-pVQZ, except for the a-g, c-g and d-g off-diagonal dipole moments where CASSCF calculations were used. As far as possible, we use SS-CASSCF calculations to obtain the orbitals for icMRCI. The $B^4\Pi$

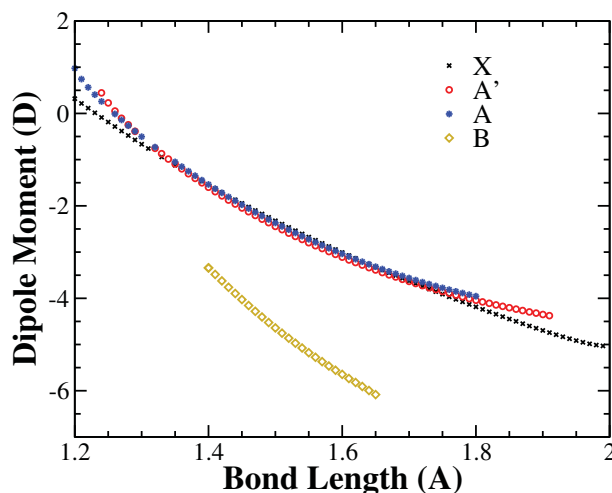


Figure 6. Diagonal dipole moment of quartet states.

state wavefunction used (A',A,B)-CAS, the $g^2\Pi$ state wavefunction used (e,f,g)-CAS and the $C^4\Sigma^-$ state used (X,C)-CAS. icMRCI(1) calculations were used for the $X^4\Sigma^-$, $A'^4\Phi$, $D^4\Delta$ (using A_2 symmetry), $a^2\Sigma^-$, $b^2\Gamma$ (using A_2 symmetry) and $e^2\Phi$ states, while icMRCI(2) calculations (requesting 2 states in the C_{2v} icMRCI) were used for the $A^4\Pi$, $c^2\Delta$ (using A_2 symmetry) and $f^2\Pi$ and icMRCI(3) calculations were used for the $d^2\Sigma^+$ state. For spin-orbit coupling elements, the CASSCF is state-averaged over the states required for the bra and ket wavefunction.

Following recent recommendations to improve reproducibility of quantum chemistry results [35], we include sample input files in the Supplementary Information.

3.2. Results and discussion

Full quantitative results are provided in the Supplementary Information in tabular format.

3.2.1. Diagonal dipole moments

Diagonal dipole moments are only needed for those states that are expected to have significant thermal population at 5000 K; we choose to include the four lowest quartet and five lowest doublet states here.

Figures 6 and 7 show the *ab initio* data for the quartet and doublet states, respectively. For all states except the $a^2\Sigma^-$ state, it is clear that the electronic state is of largely ionic character due to the near linearity of the dipole moment with increased bond length. The $X^4\Sigma^-$, $A'^4\Phi$, $A^4\Pi$, $b^2\Gamma$, $d^2\Sigma^+$ and $e^2\Phi$ states have similar charge distributions (i.e. ionic separation of charge), while the $B^4\Pi$ state has a much larger asymmetric distribution of electric charge, causing the significant increase in the magnitude of the $B^4\Pi$ state dipole moment. This is consistent with

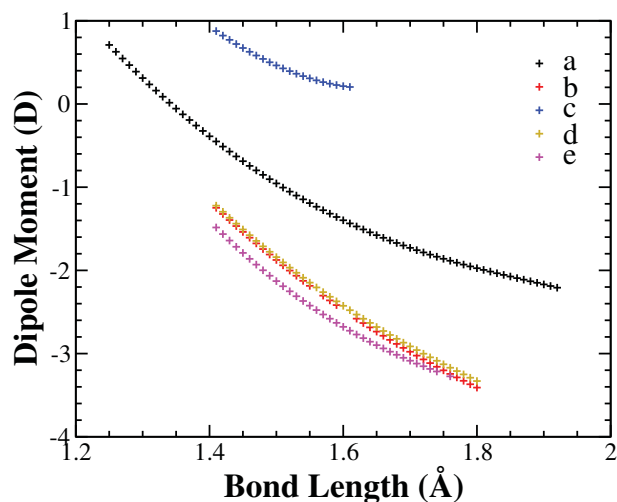


Figure 7. Diagonal dipole moment of doublet states.

the results of Ref. [14]. On a basic level, the B state has $3d^5$ valence occupancy whilst the others have $4s^1 3d^4$ valence occupancy. From this, the adverse effects of including the B $^4\Pi$ state in the CASSCF orbital optimisation for the A $^4\Pi$ state are clearly explained.

Despite the ionic characters of the VO electronic states in the region studied, these states all disassociate to neutral atomic V+O. Therefore, at some point, there is an ionic/covalent avoided crossing. This is the reason why the *ab initio* results often do not extend beyond about 1.8 Å. Convergence beyond this region was difficult because the ionic and covalent states were of similar energies. However, the dipole moment obtained is very sensitive to the ionic/covalent character of the state. Therefore, even if calculations converged, the dipole moment obtained was often unpredictable and not smooth in this region. Changes in basis set and/or method etc. did not significantly improve smoothness and convergence and so our methodology was not changed here.

Details on the way in which these diagonal dipole moments are used in the final line list are given in Ref. [80]. In brief, for our final line list, we choose to use *ab initio* points only where the calculations were trusted and the curves smooth. Using a diabatic-type approach to the ionic/covalent avoided crossing, we derived an appropriate functional form for the shape of the dipole moment going from ionic to covalent character. We then fit these *ab initio* points to the functional form to obtain smooth curves with the correct physics for our spectroscopic model.

However, this is unsatisfying from a quantum chemistry perspective. Fundamentally, the CASSCF orbitals struggle to represent both the ionic and covalent structures fairly and consistently. It would be preferable to obtain different orbitals for both states, then use both

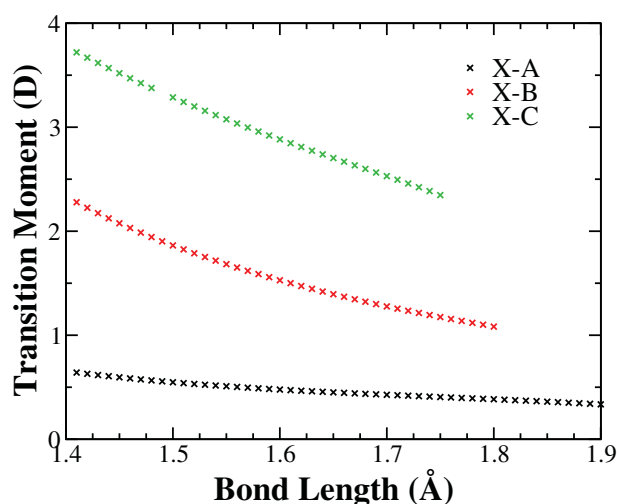


Figure 8. The main three off-diagonal dipole moments curves for the absorption spectroscopy of VO.

of these orbitals sets (that may be non-orthogonal) in a subsequent MRCI-type calculation. New multi-reference methodologies as well as approaches based on the old but under-utilised valence-bond orbitals could help address this problem.

3.2.2. Off-diagonal dipole moments

Figure 8 shows the final results for the off-diagonal dipole moment for the three main bands in VO absorption spectra. The quality of these curves is the major factor in the quality of the final line list. All three curves are smooth and tend consistently to zero at long bond distances, as required. The C-X transition moment is significantly stronger than the B-X transition moment, which is significantly stronger than the A-X transition moment. The strength of the final absorption is approximately proportional to the square of this transition moment multiplied by the frequency of the transition and so can be expected to follow this trend for a flat input light source. However, the input light is generally a black-body like; therefore, the importance of absorption in each of these three bands on the atmospheric physics depends on the temperature of the light source. Generally, VO is present in the atmospheres M dwarf stars which emit black-body radiation with temperatures around 2000 K. The emission of these objects peaks near the A-X 0-0 transition and therefore this weaker transition will probably have the most influence on the radiative transfer within the M dwarf. Conversely for hot Jupiters illuminated by hotter stars, the black-body peak will move towards the B-X and C-X transition, which will then become more important in the physics of the hot Jupiter atmosphere. All three of these transition moments should be accurate to about 5%–10%, given our considerations in the previous section.

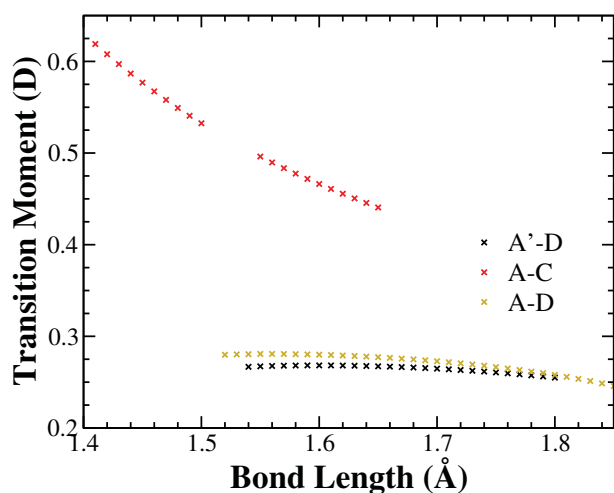
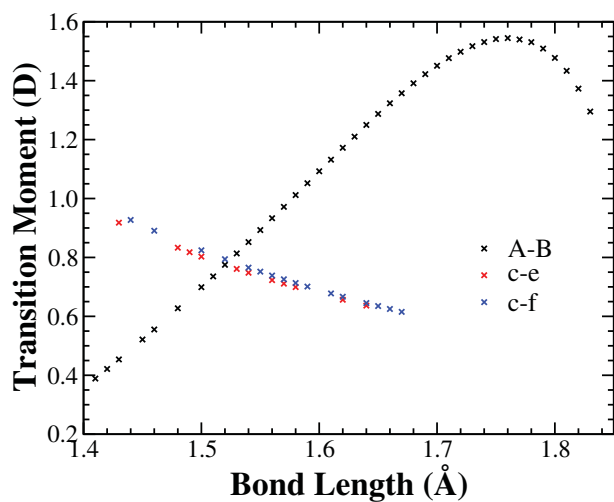


Figure 9. Off-diagonal dipole moment curves.

Figures 9 and 10 show the other off-diagonal dipole moments we consider. These are all transition moments that are not zero by symmetry or spin that both originate from electronic states with term energies less than $16,000\text{ cm}^{-1}$ (i.e. A' , A , B , a , b , c , d , e) and go to one of the lowest 13 electronic states. All curves are relatively smooth as required for our final application; achieving smoothness required careful selection of points in many cases. Most curves go smoothly from a peak towards zero. The main exceptions are the A - B , A' - D , A - D and B - C transitions. The unusual shape of the A - B spectra, however, is of some importance. Other basis sets and method choices did not significantly affect this curve, leading us to conclude that this is probably the true shape of this transition moment. The A' - D , A - D and B - C transition all have similar shapes that are flatter than would be expected; however, these are relatively weak and were not investigated further at this stage.

The availability of experimental results, particularly on excitation energies and vibrational frequencies, can

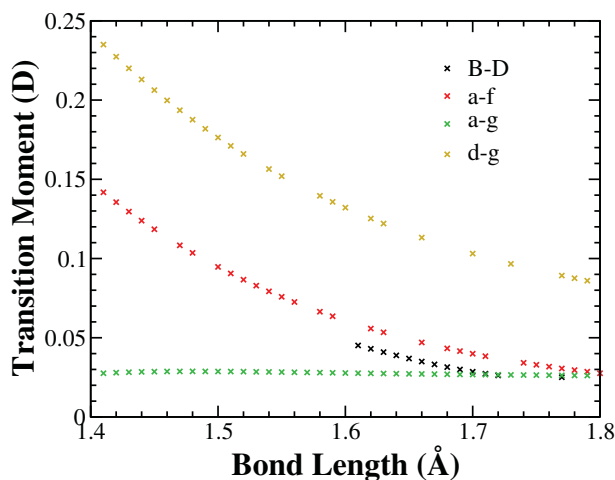
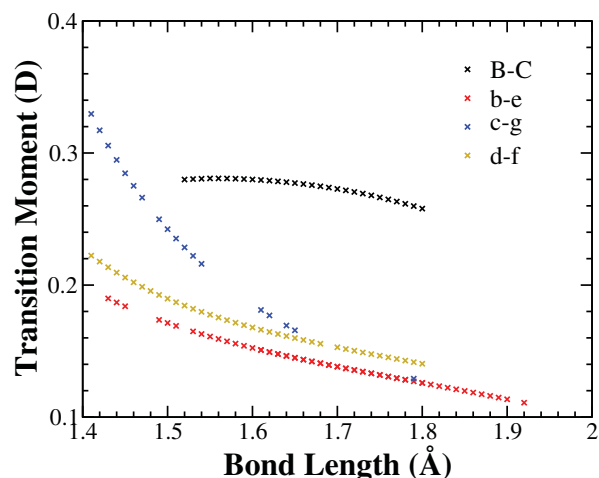


Figure 10. Off-diagonal dipole moment curves, continued.

greatly increase the accuracy of the final spectroscopic model [81] and help correct for *ab initio* inaccuracies. There are many experimental results available for VO [52–57,59,73,82–84] though more data, particularly for the higher vibrational states, would be very useful.

3.2.3. Diagonal spin-orbit coupling curves

The diagonal spin-orbit coupling curves are shown in Figure 11, with experimental equilibrium values when known given by dotted horizontal lines. The agreement between experiment and theory is generally quite high. The variation of this coupling with bond length is reasonably small. The significant differences in the magnitudes of the various SO coupling constants and the relatively high accuracy of the *ab initio* calculations mean that this property can be used as a key characteristic allowing matching of experimental and theoretical electronic states at energies where the state density is large and the absolute value and relative positioning of energy levels is not sufficiently accurate.

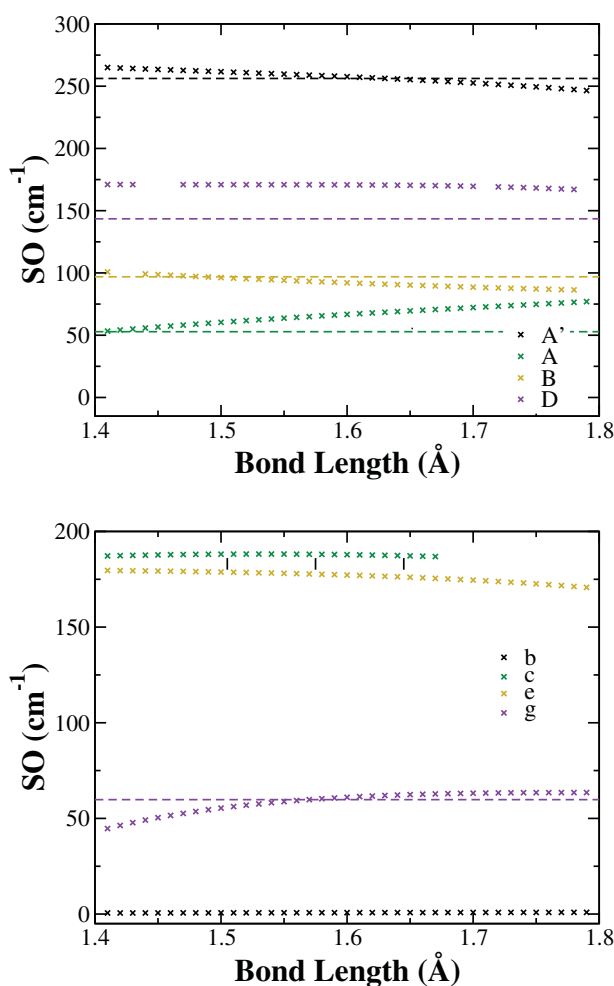


Figure 11. (Colour online) Diagonal spin–orbit coupling curves. Crosses indicate *ab initio* data points. Horizontal dotted lines indicate empirical $v = 0$ spin–orbit coupling constants from Merer [85]. The black vertical lines in the bottom plot illustrate visually the empirical difference between the $c^2\Delta$ and $e^2\Phi$ spin–orbit coupling constant (the absolute value of these constants has not been extracted from experiment).

3.2.4. Off-diagonal spin–orbit coupling curves

The off-diagonal spin–orbit coupling between states of the same spin are shown in Figure 12. The relative sizes of the couplings can be seen visually. Their effect on the energy levels will depend approximately on the square of the coupling divided by the energy gap. The a–d coupling is particularly strong; however, the influence of this coupling on the absorption spectra of VO is likely to be minimal. The f–g coupling is also quite large, yet there is no experimental evidence of this coupling. This may arise due to the MRCI methodology mixing the $f^2\Pi$ and $g^2\Pi$ states compared to the ‘true’ answer. A similar argument applies to the A–B spin–orbit coupling, but this is much smaller in magnitude.

The off-diagonal spin–orbit couplings between states of different spin are shown in Figure 13. These couplings

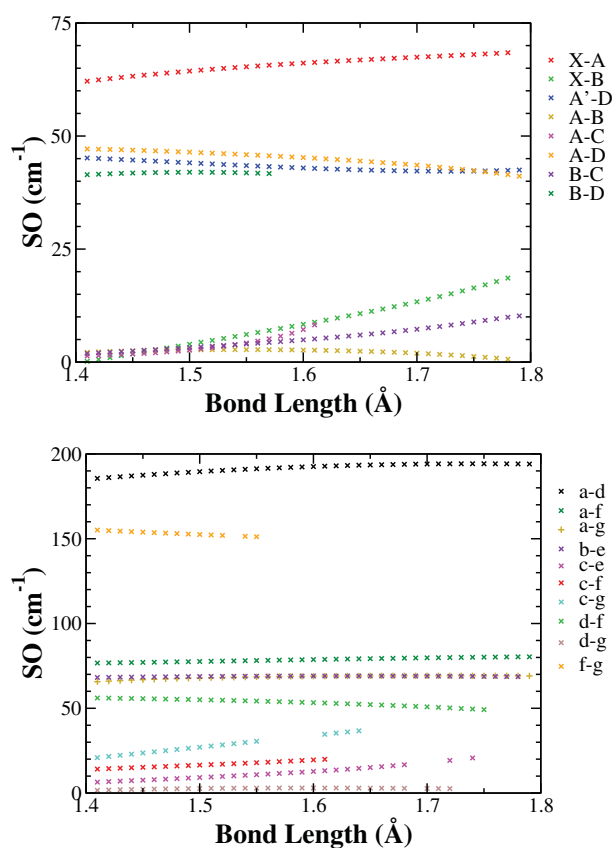


Figure 12. Off-diagonal spin–orbit coupling curves between states of the same spin.

play a very special role in being the only mechanism to allow mixing of electronic states with different spins. This means that in a VO model that incorporates these coupling terms, the final rovibronic energy levels will not be pure doublet or quartet nature. This gives rise to spin–forbidden transitions. The degree to which these coupling influence energy levels is determined by, approximately, the square of the spin–orbit coupling divided by the energy difference between the two electronic states. Therefore, it is clear that the X–d and A–g spin–orbit couplings, being over 200 cm^{-1} , can cause significant perturbations, mixing and forbidden transitions. This coupling between the X–d state could be the reason why the X–f and X–g transitions can be observed in [59]. Further, the A'–b and A–d spin–orbit couplings are likely to have a significant influence, given the proximity of the two states and the relatively large magnitude of the coupling. When considering the results for the higher electronic states, it is important to note that there are a significant number of doublet states just above the 13 electronic states considered here [15]. The coupling to these states may be as significant as, say, the D–f and D–g couplings quantified in the bottom plot of Figure 13. On the scale of Figure 13, the off-diagonal spin–orbit couplings seem relatively

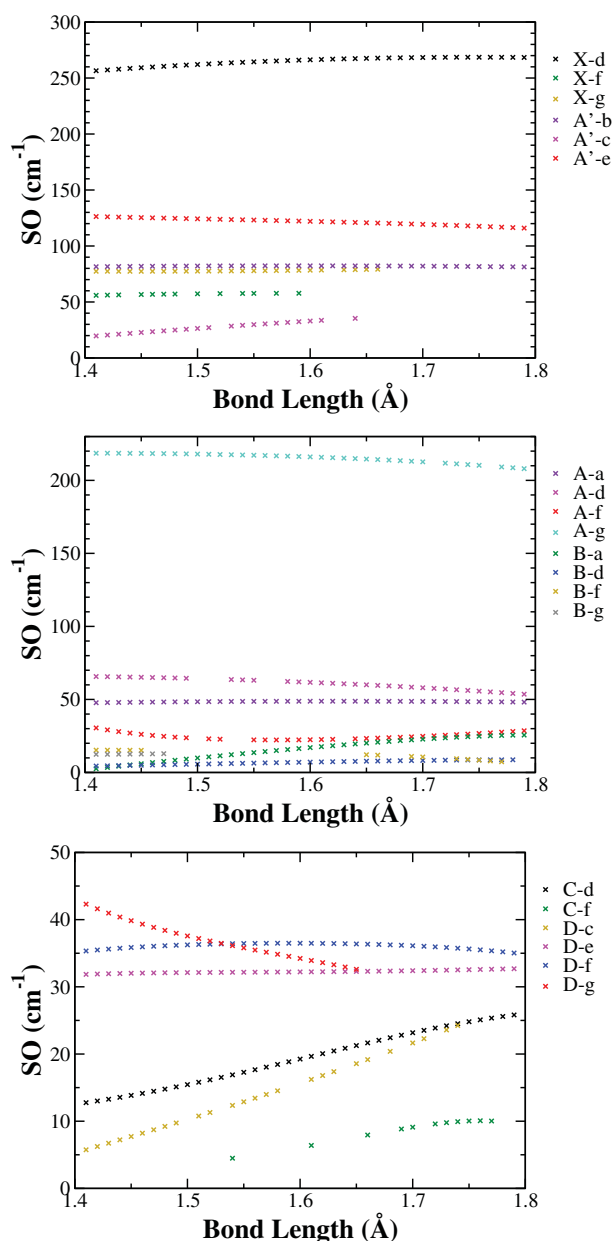


Figure 13. Off-diagonal spin–orbit coupling curves between states of the different spin.

constant across the bond length; therefore, use of an equilibrium value may be appropriate for many comparisons.

4. Conclusion

The use *ab initio* quantum chemistry for the study of electronic excited states of transition metal diatomics is definitely not at the ‘black-box’ stage. Results need to be considered carefully and critically; assessing convergence with respect to basis set, orbitals and method can be a useful way of assessing to what extent results can be trusted in the absence of preferred experimental data.

We provide some of the first results for off-diagonal dipole moments calculated using the FD methodology. Unfortunately, we find significant issues that prevent its use for transition metal diatomic systems currently. Primarily, there is significant ambiguity over which energy difference to use in the finite-field difference formula. Given the large errors in the excitation energies with current *ab initio* methodologies, this issue inhibits accuracy in the finite-field off-diagonal dipole moment. Furthermore, for perpendicular transitions, calculations times are significantly increased, by about an order of magnitude. This is due primarily to the fact that the applied electric field has to be perpendicular to the molecular axis which reduces the symmetry of the system.

These problems with evaluating accurate transition moments for transition metal diatomics fundamentally arises due to the inadequacy of orthogonal one-electron orbitals as a good first-order approximation of the electron distribution around the transition metal centre. Using a large active space helps alleviate, but does not fully solve, these issues. Differing occupancies of the *3d* and *4s* orbitals radically change the characteristics of these orbitals. Thus, ideally, different *3d* and *4s* occupancies should be represented by different *3d* and *4s* orbitals. However, orbitals optimised for each occupancies would be non-orthogonal to each other, which cannot currently be handled in main-stream quantum chemistry packages. Nevertheless, some preliminary calculations based on this reasoning have been performed. Promising results have recently been found in atomic systems using multi-configuration methods with non-orthogonal orbitals [86–88]. A similar approach was recently tested for Cr_2 by Olsen [89]; we welcome further development on this methodology and hope in particular for more accurate electronic excitation energies, which may allow finite-field difference methodology for off-diagonal dipole moments to be sufficiently reliable and robust to be useful.

Despite this gloomy picture, there are some things that are well represented by *ab initio* methodology. Spin–orbit coupling constants are often surprisingly accurate, even with low level theory, as long as the electronic state is correctly identified. The spin–orbit coupling constants are a great way to relate an experimentally observed state to a theoretical prediction, or to identify the same state in different theoretical calculations.

This work produces recommended *ab initio* data points for the spectroscopic study of VO, which are given in full in the Supplementary Information. These will be used to fit a full spectroscopic model for VO and, from this model, a full spectroscopic line list for this system. Work on this is well advanced and the results will be reported elsewhere [80].

Elsewhere [80], we report the refinement of this data to produce a final spectroscopic model for VO, the evaluation of accurate variational solutions using the newly published flexible, nuclear motion code called DUO [60], and a line list (list of rovibronic energy levels and the intensities of transitions between these levels).

Acknowledgments

We thank Lorenzo Lodi and Maire Gorman for helpful discussions on this topic. We also thank Jeffrey Reimers for his advice in manuscript preparation. This work was supported by the ERC under the Advanced Investigator Project 267219. The authors acknowledge the use of the UCL Legion High Performance Computing Facility (Legion@UCL), and associated support services, in the completion of this work.


Disclosure statement

No potential conflict of interest was reported by the authors.

Funding

ERC: the Advanced Investigator Project 267219.

ORCID

Laura K. McKemmish  <http://orcid.org/0000-0003-1039-2143>

References

- [1] M. Castelaz, D. Luttermoser, and R. Piontek, *Astrophys. J.* **538**(1, Part 1), 341 (2000).
- [2] H. Hoeijmakers, R. de Kok, I. Snellen, M. Brogi, J. Birkby, and H. Schwarz, *Astron. Astrophys.* **575**, A20 (2015).
- [3] D. Pudasainee, H.R. Paur, S. Fleck, and H. Seifert, *Fuel Process. Technol.* **120**, 54 (2014).
- [4] P. Monkhouse, *Prog. Energ. Combust.* **28**, 331 (2002).
- [5] C. Bauschlicher and P. Maitre, *Theor. Chim. Acta* **90**(2–3), 189 (1995).
- [6] M. Pykavy and C. van Wullen, *J. Phys. Chem. A* **107**(29), 5566 (2003).
- [7] B. Dai, K. Deng, J. Yang, and Q. Zhu, *J. Chem. Phys.* **118**(21), 9608 (2003).
- [8] E. Broclawik and T. Borowski, *Chem. Phys. Lett.* **339**, 433 (2001).
- [9] A. Bridgeman and J. Rothery, *J. Chem. Soc. Dalton* (2), 211 (2000).
- [10] M. Calatayud, B. Silvi, J. Andres, and A. Beltran, *Chem. Phys. Lett.* **333**(6), 493 (2001).
- [11] D. Quan, W. Ling, S. Xiao-Hong, W. Hong-Yan, G. Tao, and Z. Zheng-He, *Acta Chimica Sinica* **66**(1), 23 (2008).
- [12] H.J. Kulik and N. Marzari, *J. Chem. Phys.* **133**(11) (2010).
- [13] D. Quan, W. Ling, S. Xiao-Hong, and G. Tao, *Acta Physica Sinica* **55**(12), 6308 (2006).
- [14] E. Miliordos and A. Mavridis, *J. Phys. Chem. A* **111**(10), 1953 (2007).
- [15] O. Hübner, J. Hornung, and H.J. Himmel, *J. Chem. Phys.* **143**(2), 024309 (2015).
- [16] J.F. Harrison, *Chem. Rev.* **100**, 679 (2000).
- [17] A. Kalemios and A. Mavridis, *J. Chem. Phys.* **113**, 2270 (2000).
- [18] A.C. Borin and L.G.M. de Macedo, *Chem. Phys. Lett.* **383**, 53 (2004).
- [19] D. Tzeli and A. Mavridis, *J. Chem. Phys.* **118**(11), 4984 (2003).
- [20] S.R. Langhoff, *Astrophys. J.* **481**, 1007 (1997).
- [21] N.V. Dobrodey, *Astron. Astrophys.* **365**, 642 (2001).
- [22] K.C. Namiki, H. Saitoh, and H. Ito, *J. Mol. Spectrosc.* **226**, 87 (2004).
- [23] E. Miliordos and A. Mavridis, *J. Phys. Chem. A* **114**, 8536 (2010).
- [24] N.J. DeYonker, D.T. Halfen, W.D. Allen, and L.M. Ziurys, *J. Chem. Phys.* **141**, 204302 (2014).
- [25] E. Miliordos and K.L.C. Hunt, *J. Phys. Chem. A* **115**, 4436 (2011).
- [26] C.N. Sakellaris and A. Mavridis, *J. Chem. Phys.* **138**(5), 054308 (2013).
- [27] C.N. Sakellaris and A. Mavridis, *J. Chem. Phys.* **137**(3), 034309 (2012).
- [28] C.N. Sakellaris, A. Papakondylis, and A. Mavridis, *J. Phys. Chem. A* **114**(34), 9333 (2010).
- [29] C.N. Sakellaris, E. Miliordos, and A. Mavridis, *J. Chem. Phys.* **134**, 234308 (2011).
- [30] S. Koseki, Y. Ishihara, H. Umeda, D.G. Fedorov, and M.S. Gordon, *J. Phys. Chem. A* **106**, 785 (2002).
- [31] L. Lodi, S.N. Yurchenko, and J. Tennyson, *Mol. Phys.* **113**(13), 1998 (2015).
- [32] S. Koseki, Y. Ishihara, D.G. Fedorov, H. Umeda, M.W. Schmidt, and M.S. Gordon, *J. Phys. Chem. A* **108**, 4707 (2004).
- [33] K. Tanaka, M. Sekiya, and M. Yoshimine, *J. Chem. Phys.* **115**, 4558 (2001).
- [34] S. Koseki, T. Matsushita, and M. Gordon, *J. Phys. Chem. A* **110**, 2560 (2006).
- [35] J. Tennyson, L. Lodi, L.K. McKemmish, and S.N. Yurchenko, *J. Phys. B: At. Mol. Opt. Phys.* **49**, 102001 (2016).
- [36] U. Schollwöck, *Rev. Mod. Phys.* **77**(1), 259 (2005).
- [37] K.H. Marti, I.M. Ondík, G. Moritz, and M. Reiher, *J. Chem. Phys.* **128**(1), 014104 (2008).
- [38] Y. Kurashige and T. Yanai, *J. Chem. Phys.* **135**(9), 094104 (2011).
- [39] G.K.L. Chan and S. Sharma, *Annu. Rev. Phys. Chem.* **62**, 465 (2011).
- [40] T. Yanai, Y. Kurashige, W. Mizukami, J. Chalupská³, T.N. Lan, and M. Saitow, *Intern. J. Quantum Chem.* **115**(5), 283 (2015).
- [41] J. Coe, P. Murphy, and M. Paterson, *Chem. Phys. Lett.* **604**, 46 (2014).
- [42] R.E. Thomas, Q. Sun, A. Alavi, and G.H. Booth, *J. Chem. Theo. Comp.* **11**(11), 5316 (2015).
- [43] P. Pulay, *Int. J. Quantum Chem.* **111**(13), 3273 (2011).
- [44] H.K. Chung, B.J. Braams, K. Bartschat, A.G. Császár, G.W.F. Drake, T. Kirchner, V. Kokoouline, and J. Tennyson, *J. Phys. D: Appl. Phys.* **49**(36), 363002 (2016).
- [45] H. Werner and P.J. Knowles, *J. Chem. Phys.* **82**, 5053 (1985).

- [46] H.J. Werner, P.J. Knowles, G. Knizia, F.R. Manby, and M. Schütz, *WIREs Comput. Mol. Sci.* **2**, 242 (2012).
- [47] H. Werner and P.J. Knowles, *J. Chem. Phys.* **89**, 5803 (1988).
- [48] J. Tennyson, *J. Mol. Spectrosc.* **298**, 1 (2014).
- [49] L. Lodi, S.N. Yurchenko, and J. Tennyson, *Mol. Phys.* **113**, 1559 (2015).
- [50] A.P.L. Rendell, G.B. Bacskay, N.S. Hush, and N.C. Handy, *J. Chem. Phys.* **87**, 5976 (1987).
- [51] S.O. Adamson, A. Zaitsevskii, and N.F. Stepanov, *J. Phys. B: At. Mol. Opt. Phys.* **31**, 5275 (1998).
- [52] A. Cheung, R. Hansen, and A. Merer, *J. Mol. Spectrosc.* **91**(1), 165 (1982).
- [53] A. Cheung, A. Taylor, and A. Merer, *J. Mol. Spectrosc.* **92**(2), 391 (1982).
- [54] A. Merer, G. Huang, A. Cheung, and A. Taylor, *J. Mol. Spectrosc.* **125**, 465 (1987).
- [55] R. Suenram, G. Fraser, F. Lovas, and C. Gillies, *J. Mol. Spectrosc.* **148**(1), 114 (1991).
- [56] A. Cheung, P. Hajigeorgiou, G. Huang, S. Huang, and A. Merer, *J. Mol. Spectrosc.* **163**(2), 443 (1994).
- [57] A.G. Adam, M. Barnes, B. Berno, R.D. Bower, and A.J. Merer, *J. Mol. Spectrosc.* **170**, 94 (1995).
- [58] M. Flory and L. Ziurys, *J. Mol. Spectrosc.* **247**, 76 (2008).
- [59] W.S. Hopkins, S.M. Hamilton, and S.R. Mackenzie, *J. Chem. Phys.* **130**(14) (2009).
- [60] S.N. Yurchenko, L. Lodi, J. Tennyson, and A.V. Stolyarov, *Comput. Phys. Commun.* **202**, 262 (2016).
- [61] S.R. Langhoff and E.R. Davidson, *Intern. J. Quantum Chem.* **8**(1), 61 (1974).
- [62] C.W. Bauschlicher, S.R. Langhoff, P.R. Taylor, N.C. Handy, and P.J. Knowles, *J. Chem. Phys.* **85**(3), 1469 (1986).
- [63] C.W. Bauschlicher and S.R. Langhoff, *J. Chem. Phys.* **89**, 2116 (1988).
- [64] P. Slavíček and T.J. Martínez, *J. Chem. Phys.* **132**, 234102 (2010).
- [65] A.S. de Meras, M.B. Lepetit, and J.P. Malrieu, *Chem. Phys. Lett.* **172**, 163 (1990).
- [66] A. Zaitsevskii and J.P. Malrieu, *Chem. Phys. Lett.* **228**, 458 (1994).
- [67] N. Guihery, J.P. Malrieu, D. Maynau, and K. Handrick, *Intern. J. Quantum Chem.* **61**, 45 (1997).
- [68] C. Angeli, C.J. Calzado, R. Cimraglia, S. Evangelisti, and D. Mayna, *Mol. Phys.* **101**, 1937 (2003).
- [69] M.W. Schmidt and M.S. Gordon, *Annu. Rev. Phys. Chem.* **49**, 233 (1998).
- [70] K.K. Docken and J. Hinze, *J. Chem. Phys.* **57**(11) (1972).
- [71] Y. Suzuki, S. Asai, K. Kobayashi, T. Noro, F. Sasaki, and H. Tatewaki, *Chem. Phys. Lett.* **268**(3–4), 213 (1997).
- [72] O. Krechkivska, G.B. Bacskay, T.P. Troy, K. Nauta, T.D. Kreuscher, S.H. Kable, and T.W. Schmidt, *J. Phys. Chem. A* **119**(50), 12102 (2015).
- [73] L. Karlsson, B. Lindgren, C. Lundevall, and U. Sassenberg, *J. Mol. Spectrosc.* **181**(2), 274 (1997).
- [74] J. Tennyson, K. Hulme, O.K. Naim, and S.N. Yurchenko, *J. Phys. B: At. Mol. Opt. Phys.* **49**, 044002 (2016).
- [75] G.H. Dierksen, B.O. Roos, and A.J. Sadlej, *Chem. Phys.* **59**, 29 (1981).
- [76] M. Ernzerhof, C.M. Marian, and S.D. Peyerimhoff, *Intern. J. Quantum Chem.* **43**, 659 (1992).
- [77] J. Lipiński, *Chem. Phys. Lett.* **363**, 313 (2002).
- [78] L. Lodi and J. Tennyson, *J. Phys. B: At. Mol. Opt. Phys.* **43**, 133001 (2010).
- [79] J. Stanton, J. Gauss, M. Harding, P. Szalay, A. Auer, R. Bartlett, U. Benedikt, C. Berger, D. Bernholdt, and Y. Bomble, CFour, For the current version, see <http://www.cfour.de> (2009).
- [80] L.K. McKemmish, S.N. Yurchenko, and J. Tennyson, *Mon. Not. R. Astron. Soc.* accepted (2016).
- [81] A.T. Patrascu, C. Hill, J. Tennyson, and S.N. Yurchenko, *J. Chem. Phys.* **141**, 144312 (2014).
- [82] R.S. Ram, P.F. Bernath, S.P. Davis, and A.J. Merer, *J. Mol. Spectrosc.* **211**, 279 (2002).
- [83] R.S. Ram and P.F. Bernath, *J. Mol. Spectrosc.* **229**, 57 (2005).
- [84] M.A. Flory and L.M. Ziurys, *J. Mol. Spectrosc.* **247**(1), 76 (2008).
- [85] A. Merer, *Annu. Rev. Phys. Chem.* **40**, 407 (1989).
- [86] J. Olsen, M.R. Godefroid, P. Jonsson, P.A. Malmqvist, and C.F. Fischer, *Phys. Rev. E* **52**, 4499 (1995).
- [87] A. Toner and A. Hibbert, *Mon. Not. R. Astron. Soc.* **361**(2), 673 (2005).
- [88] S.S. Tayal, *At. Data and Nucl. Data Tables* **97**, 481 (2011).
- [89] J. Olsen, *J. Chem. Phys.* **143**, 114102 (2015).