

Molecular Physics

An International Journal at the Interface Between Chemistry and Physics

ISSN: 0026-8976 (Print) 1362-3028 (Online) Journal homepage: <https://www.tandfonline.com/loi/tmph20>

A global potential energy surface for H_3^+

Irina I. Mizus, Oleg L. Polyansky, Laura K. McKemmish, Jonathan Tennyson, Alexander Alijah & Nikolai F. Zobov

To cite this article: Irina I. Mizus, Oleg L. Polyansky, Laura K. McKemmish, Jonathan Tennyson, Alexander Alijah & Nikolai F. Zobov (2019) A global potential energy surface for H_3^+ , Molecular Physics, 117:13, 1663-1672, DOI: [10.1080/00268976.2018.1554195](https://doi.org/10.1080/00268976.2018.1554195)

To link to this article: <https://doi.org/10.1080/00268976.2018.1554195>



Published online: 05 Dec 2018.



Submit your article to this journal [↗](#)



Article views: 75



View Crossmark data [↗](#)



Citing articles: 1 View citing articles [↗](#)

A global potential energy surface for H_3^+

Irina I. Mizus^a, Oleg L. Polyansky^{a,b}, Laura K. McKemmish^{b,c}, Jonathan Tennyson^b, Alexander Alijah^{b,d} and Nikolai F. Zobov^a

^aInstitute of Applied Physics, Russian Academy of Sciences, Nizhny Novgorod, Russia; ^bDepartment of Physics and Astronomy, University College London, London, UK; ^cDepartment of Chemistry, University of New South Wales, Sydney, Australia; ^dGroupe de Spectrométrie Moléculaire et Atmosphérique, GSMA, UMR CNRS 7331, Université de Reims Champagne-Ardenne, Reims, France

ABSTRACT

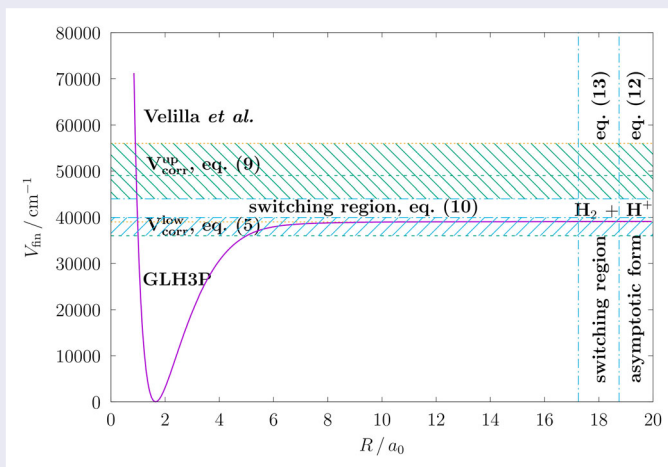
A globally correct potential energy surface (PES) for the H_3^+ molecular ion is presented. The Born-Oppenheimer (BO) *ab initio* grid points of Pavanello *et al.* [*J. Chem. Phys.* **136**, 184303 (2012)] are refitted as BOPE575K, which reproduces the energies below dissociation with a root mean square deviation of 0.05 cm^{-1} ; points between dissociation and $75,000 \text{ cm}^{-1}$ are reproduced with the average accuracy of a few wavenumbers. The new PES75K+ potential combines BOPE575K with adiabatic, relativistic and quantum electrodynamics (QED) surfaces to provide the most accurate representation of the H_3^+ global potential to date, overcoming the limitations on previous high accuracy H_3^+ PESs near and above dissociation. PES75K+ can be used to provide predictions of bound rovibrational energy levels with an accuracy of approaching 0.1 cm^{-1} . Calculation of rovibrational energy levels within PES75K+ suggests that the non-adiabatic correction remains a limiting factor. The PES is also constructed to give the correct asymptotic limit making it suitable for use in studies of the $\text{H}^+ + \text{H}_2$ prototypical chemical reaction. An improved dissociation energy for H_3^+ is derived as $D_0 = 35,076 \pm 2 \text{ cm}^{-1}$.

ARTICLE HISTORY

Received 24 September 2018
Accepted 20 November 2018

KEYWORDS

Ab initio; potential energy surface; dissociation; spectroscopy



1. Introduction

The H_3^+ ion provides a benchmark system for two areas of science, which, up to now, have remained unrelated: the high accuracy, *ab initio* prediction of rotation-vibration spectra [1] and reaction dynamics [2]. In fact, the two regimes are linked through the near-dissociation spectrum of Carrington and co-workers

[3–8], which provides a direct connection between spectroscopy and dissociation dynamics. Theoretical studies to elucidate this spectrum [9], as well as studies, which try to model ultra-low energy $\text{H}^+ + \text{H}_2$ reactive and non-reactive scattering [10–12], require surfaces of spectroscopic accuracy to recover the full resonance structure.

A number of global potential energy surfaces (PESs) are available for H_3^+ in its electronic ground state [13–23]. Of these we particularly note the GLH3P PES of Pavanello *et al.* [21], which is based on *ab initio* calculations of spectroscopic accuracy [24, 25], and the surface of Velilla *et al.* [18], which is based on less accurate *ab initio* calculations but whose full treatment of the long-range proton – H_2 interactions is vital for the study of ultra-low energy collisions [12].

The *ab initio* calculations of Pavanello *et al.* [21] used optimised explicitly correlated shifted Gaussian functions and were performed for an extensive grid of 41,655 H_3^+ geometries. The calculations have an absolute accuracy of about 0.01 cm^{-1} for the non-relativistic, fixed-nuclei, electronic energy [26]; Pavanello *et al.*'s analytical fit reproduces these points below dissociation with a root mean square error of about 0.05 cm^{-1} . However, certain asymptotic configurations were omitted from this fit as their inclusion led to significant deterioration in the accuracy at low energies. Therefore, these points are not well represented by the analytical GLH3P yielding unphysical features as shown in the supplemental material of the original paper [21]. Such unphysical features also appear in other global H_3^+ PESs [15]. Importantly, we note that the PES of Velilla *et al.* [18] does not show any unphysical behaviour.

The work of Pavanello *et al.* [21] also included an adiabatic or diagonal Born-Oppenheimer correction, but did not provide a global fitted surface for this effect. However, to make a PES spectroscopically accurate, global

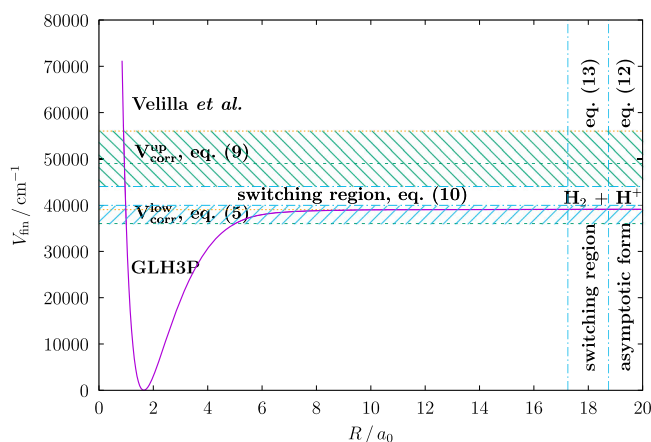


Figure 1. Illustration of the structure of new BOPES75K potential energy surface. The curve corresponds to isosceles configurations of H_3^+ with the third bond length fixed at its equilibrium value $1.65 a_0$. Dashed lines mark the borders of regions, where the GLH3P PES (the green line) and the potential of Velilla *et al.* (the yellow line) were corrected, and the blue dashed lines denote the borders of switching regions. The shaded domains correspond to the parts of corrected potentials, which do not fall into the energy switching region.

adiabatic, relativistic and QED (quantum electrodynamics) correction surfaces should also be included. Such surfaces are available [1]. In the present work, we report a significantly improved fit to the *ab initio* points of Pavanello *et al.* [21] in the high-energy region, so that a very accurate global PES is obtained, which includes BO energies, QED, relativistic and adiabatic corrections. This surface, which we call PES75K+, is suitable for both the computation of predissociation energies and line positions, as well as for accurate reactive and non-reactive scattering calculations.

The present paper is organised as follows: In Section 2 we describe the functional form of the improved global H_3^+ PES. The global PES obtained as an analytical form fitted to the accurate *ab initio* points is analyzed in Section 3, while Section 4 contains a comparison of the rovibrational energy level calculations using the present surface and our previous one. Section 5 concludes our paper.

2. Structure of the new global potential

2.1. Born-Oppenheimer part of the potential

Initially we attempted to fit the *ab initio* data of Pavanello *et al.* [21] using a single functional form. However, our attempts failed to give a satisfactory fit. We therefore adopted an alternative approach where the BO part of the new *ab initio* high-accuracy global potential energy surface for the H_3^+ molecule consists of three independent parts: low and high energy short-range potentials joined together smoothly at $42,000 \text{ cm}^{-1}$ using the energy switching approach of Varandas [27], and an analytic long-range part, which gives the correct asymptotic behaviour. The lower part of the short-range potential, V_{low} , consists of the three-body terms from the GLH3P potential obtained by Pavanello *et al.* [21] with a correction in the energy region where the GLH3P surface is not accurate. The upper part, V_{up} , is an analogous adaptation of the PES of Velilla *et al.* [18]. The long-range potential is based on the analytic form used by Velilla *et al.* [18]. We call the resulting PES BOPES75K. Figure 1 illustrates the structure of BOPES75K.

The BOPES75K potential extensively utilises switching functions,

$$f(t - t_0, \alpha) = \frac{1}{1 + e^{-2\alpha\Delta t}} = \frac{1}{2} [1 + \tanh(\alpha\Delta t)], \quad (1)$$

where α is a parameter that controls the sharpness of the switch (large α is a fast switch and small α is a slow switch), t_0 is the switching point in variable t and $\Delta t = t - t_0$ is the distance from the position of the switch. However, it is often easier to set a width for the switching

zone, b , so that at the edges of the region ($t_0 - b/2, t_0 + b/2$) the switching function would reach its asymptotic values (0 or 1, correspondingly) with an accuracy of δ :

$$f\left(t_0 + \frac{b}{2}\right) = 1 - \frac{\delta}{100},$$

$$f\left(t_0 - \frac{b}{2}\right) = \frac{\delta}{100},$$

where δ is expressed as a percentage. So, the sharpness of the switch is fully defined by the width of the switching zone and the accuracy δ and can be expressed as

$$\alpha = \ln\left(\frac{100}{\delta} - 1\right) / b. \quad (2)$$

We will use the parameters t_0 , b and δ later as variables in some optimisation procedures.

A product of two switching functions can be used to smoothly turn on and off a term within a finite region by using oppositely signed switching parameters. A similar approach with multiple switchings was applied to the NO_2 molecule by Varandas in [28]. In this work we use switching functions based directly on the coordinates and on the value of the potential energy at a given geometry. In the latter case it is necessary to map from the given coordinates to a value of the PES at that point in order to evaluate the switching function. Below V_d denotes the value of this ‘distributing’ potential; where necessary the global H_3^+ potential of Velilla *et al.* [18] was used to evaluate V_d .

The correction to V_{low} was made by first approximating the differences between the GLH3P energy values and the corresponding *ab initio* energy values at 2414 *ab initio* geometries in the range 35,000 to 42,000 cm^{-1} using a polynomial

$$V_{\text{poly}}^X = \sum_{nmk} V_{nmk}^X S_a^n S_e^{2m+3k} \cos(3k\varphi),$$

$$n + 2m + 3k \leq N, \quad (3)$$

where the superscript X on V_{poly}^X and V_{nmk}^X is used to distinguish the various expansions, in this case $X = \text{low}$ for the low-energy short-range PES and $N = 31$ coefficients with a maximum degree of 7. This expression uses the following symmetrised coordinates:

$$S_a = (\Delta R_{12} + \Delta R_{23} + \Delta R_{31}) / \sqrt{3},$$

$$S_x = (2\Delta R_{12} - \Delta R_{23} - \Delta R_{31}) / \sqrt{6} = S_e \cos \varphi, \quad (4)$$

$$S_y = (\Delta R_{23} - \Delta R_{31}) / \sqrt{2} = S_e \sin \varphi,$$

where $\Delta R_{ik} = R_{ik} - R_e$ is the displacement from the equilibrium value of $R_e = 1.65 a_0$ in the bond length

between the i th and k th protons in H_3^+ and the angle φ from Equation (3) can be obtained from the latter equations as $\varphi = \arctan(S_y/S_x)$. This analytical form reflects correctly the D_{3h} symmetry of the H_3^+ molecular system. The standard deviation of the approximated energy differences from the numerical values of the polynomial $V_{\text{poly}}^{\text{low}}$ in the corresponding *ab initio* geometries is about 2.3 cm^{-1} .

We added the possibility of smoothly switching on and off this polynomial term by multiplying it with switching factors f_{low} and f_{up} :

$$V_{\text{corr}}^{\text{low}} = V_{\text{poly}}^{\text{low}} f_{\text{low}} f_{\text{up}}, \quad (5)$$

where

$$f_{\text{low}} = \frac{1}{1 + e^{-2\alpha_l(V_d - E_l)}},$$

$$f_{\text{up}} = \frac{1}{1 + e^{2\alpha_u(V_d - E_u)}}, \quad (6)$$

$$\alpha_{l,u} = \ln\left(\frac{100}{\delta} - 1\right) / b_{l,u}.$$

Here and elsewhere $\delta = 0.1\%$, and b_u, b_l, E_u and E_l are the adjustable nonlinear parameters of the fit. The lower part of the potential is thus represented as

$$V_{\text{low}} = V_{\text{GLH3P}} - V_{\text{corr}}^{\text{low}}. \quad (7)$$

The upper part of our new global H_3^+ BO PES is based on the potential of Velilla *et al.* [18], denoted V_{Vel} , as the GLH3P potential is not reliable in this energy region. It is corrected in the same way as the lower part:

$$V_{\text{up}} = V_{\text{Vel}} - V_{\text{corr}}^{\text{up}}. \quad (8)$$

The correction term is also a polynomial $V_{\text{poly}}^{\text{up}}$ given by Equation (3) with $N = 30$ coefficients V_{nmk}^{up} and a maximum degree of 7. This function approximates a set of 4582 differences between the V_{Vel} values and the accurate *ab initio* energies in the region from 40,000 cm^{-1} to 55,000 cm^{-1} with a standard deviation of about 4.6 cm^{-1} . Again the function switches on and off

$$V_{\text{corr}}^{\text{up}} = V_{\text{poly}}^{\text{up}} f_{\text{low}} f_{\text{up}}, \quad (9)$$

where f_{low} and f_{up} are also given by Equation (6), just like the ones used to describe $V_{\text{corr}}^{\text{low}}$ but with their own values of adjustable nonlinear parameters.

Table 1. Adjustable nonlinear parameters used to define our new global *ab initio* PES for the H_3^+ molecule together with the number N_{coef} of linear polynomial parameters; all polynomials have a maximum degree of 7.

Part of the PES	N_{coef}	δ	b_u/cm^{-1}	b_l/cm^{-1}	E_u/cm^{-1}	E_l/cm^{-1}	E_0/cm^{-1}	g_0	g_1	R_0/a_0	b/a_0
Lower (GLH3P-like)	31	0.1	300	100	48,000	37,000	42,000	1.71	2000	18	1
Upper (Velilla <i>et al.</i> -like)	30	0.1	1400	80	55,000	40,000					

The two parts were then merged using the energy switching scheme [27] to yield a composite potential

$$V_{\text{comp}} = \varphi_{\text{low}} V_{\text{low}} + \varphi_{\text{up}} V_{\text{up}}, \quad (10)$$

where

$$\begin{aligned} \varphi_{\text{low}} &= \frac{1}{2}(1 + \tanh(-\gamma(V_{\text{up}} - E_0))), \\ \varphi_{\text{up}} &= \frac{1}{2}(1 + \tanh(\gamma(V_{\text{up}} - E_0))), \\ \gamma &= \tilde{g}_0 + \tilde{g}_1(V_{\text{up}} - E_0)^2, \end{aligned} \quad (11)$$

and the coefficients $\tilde{g}_{0,1} = g_{0,1}/(100k^2)$, where $k = 219474.631 \text{ cm}^{-1} E_h^{-1}$ is a well-known conversion factor of energy from E_h to cm^{-1} units.

The leading term in the long-range behaviour of the H_3^+ potential at its first dissociation limit into $\text{H}_2 + \text{H}^+$ is correctly represented in the GLH3P potential, as it is derived from the analytical potentials of Viegas *et al.* [17], and therefore also in V_{comp} . However, the full, angularly-dependent asymptotic behaviour is given by the multipole expansion [18, 29]

$$\begin{aligned} V - D_e \underset{R \rightarrow \infty}{\approx} & \frac{Q_2(r)P_2(\cos\theta)}{R^3} \\ & - \frac{\frac{1}{2}a_0(r) + \frac{1}{3}[\alpha_{\parallel}(r) - \alpha_{\perp}(r)]P_2(\cos\theta)}{R^4} \\ & + \dots + V_{\text{H}_2}(r), \end{aligned} \quad (12)$$

where r is the internuclear separation of the diatomic (taken as the shortest distance between two of the nuclei), R is the distance between the midpoint between these two nuclei and the third nucleus, and θ is the angle between r and R . In the case where two nearest nuclei are the same, these coordinates are standard Jacobi coordinates. We enforce this behaviour in the asymptotic region by explicitly joining our BO surface defined in Equation (10) to the one due to Velilla *et al.*, where this form is implemented. This is done at a certain distance R_0 using the switching procedure

$$V_{\text{fin}} = f_{\text{short}} V_{\text{comp}} + f_{\text{long}} V_{\text{Vel}}, \quad (13)$$

where

$$f_{\text{long}} = \frac{1}{1 + e^{-2\alpha(R-R_0)}}, \quad (14)$$

$$f_{\text{short}} = \frac{1}{1 + e^{2\alpha(R-R_0)}} \quad (15)$$

Table 2. Coefficients of the polynomials $V_{\text{poly}}^{\text{low}}$ and $V_{\text{poly}}^{\text{up}}$ given by Equation (3) for the correction terms to the lower and the upper parts of the new BOPES75K potential, correspondingly.

n	$2m$	$3k$	V_{nmk}^{low}	V_{nmk}^{up}
0	0	0	-0.75143071	28.89617715
1	0	0	-1.76440106	-10.58157805
2	0	0	-0.09859390	2.87604608
0	2	0	0.99200748	2.08628333
3	0	0	1.48567895	-0.62636049
1	2	0	-4.46772673	-0.51256565
0	0	3	-3.90622308	-1.15962111
4	0	0	-0.37926323	0.13476748
2	2	0	1.53719611	-0.08342574
1	0	3	2.48855132	0.29409500
0	4	0	1.37657184	0.10614684
5	0	0	-0.01544350	-0.02831679
3	2	0	0.07169542	0.08313826
2	0	3	-0.37054284	-0.01860622
1	4	0	-0.68993265	-0.10208867
0	2	3	-0.20471568	-0.03815698
6	0	0	0.00037056	0.00295364
4	2	0	0.00167004	-0.01213435
3	0	3	0.00764265	0.00056655
2	4	0	-0.00024218	0.01681600
1	2	3	0.02173168	0.00053141
0	6	0	0.04686634	-0.00595500
0	0	6	-0.00943583	-0.00116108
7	0	0	0.00077371	-0.00010578
5	2	0	-0.00681003	0.00050616
4	0	3	-0.00285356	-0.00000479
3	4	0	0.01863322	-0.00079326
2	2	3	0.01297023	0.00002889
1	6	0	-0.01472313	0.00040228
1	0	6	0.00114263	0.00004892
0	4	3	-0.01094351	

and α is given by Equation (2). R_0 and b are two further nonlinear parameters in our final BO PES V_{fin} , and δ has the same value as earlier, see Equation (6). This switching also ensures that the H_2 diatomic potential is correctly reproduced by the V_{fin} potential as $R \rightarrow \infty$.

Thus, the final form of our global BO PES contains 61 linear and 15 nonlinear adjustable parameters of the fit, including the parameters E_0, g_0 and g_1 for the energy switching [27]. The linear parameters were determined by least-squares fitting; the nonlinear ones were adjusted manually by a trial-and-error procedure. Their final values are summarised in Tables 1 and 2.

2.2. Correction surfaces

It is necessary to also take into account adiabatic, relativistic and QED corrections to the BO approximation. The final form of our potential PES75K+ was

obtained by addition of these correction surfaces to the BOPES75K PES.

All correction surfaces consist of two parts: a polynomial given in the form of Equation (3), and an exponential damping function, which prevents unphysical behaviour of the corrections for geometries with large internuclear distances.

2.2.1. Relativistic and QED correction surfaces

The polynomial parts $V_{\text{poly}}^{\text{rel,QED}}$ of relativistic and QED correction surfaces have the analytical form of Equation (3), but different sets of adjustable parameters: $N=52$ for the relativistic correction, which correspond to maximum polynomial power of 10, and $N=44$ coefficients V_{nmk}^{QED} for the QED correction, with maximum polynomial power of 9. These sets of parameters were obtained from two fits at points with energies below $38,000 \text{ cm}^{-1}$: (i) a fit of 3380 relativistic points by Bachorz *et al.* [30] to $V_{\text{poly}}^{\text{rel}}$, which are reproduced with a standard deviation (rms) value of 0.008 cm^{-1} , and (ii) a fit of 6413 QED points by Lodi *et al.* [31] to $V_{\text{poly}}^{\text{QED}}$ with the rms deviation of 0.001 cm^{-1} .

Our estimates show that the overall effect of relativistic and QED corrections on the H_3^+ energy states even in the near-dissociation region is only about 0.1 cm^{-1} , which is much smaller than the corresponding BOPES75K accuracy of a few cm^{-1} , and we can neglect this effect for states with large internuclear distances in the region of the first dissociation limit and above.

We included the relativistic and the QED correction surfaces using a combined polynomial form $V_{\text{poly}}^{\text{rel+QED}}$ given by Equation (3), with maximum power of 10 and polynomial coefficients $V_{nmk}^{\text{rel+QED}} = V_{nmk}^{\text{rel}} + V_{nmk}^{\text{QED}}$. The combined polynomial correction $V_{\text{poly}}^{\text{rel+QED}} = V_{\text{poly}}^{\text{rel}} + V_{\text{poly}}^{\text{QED}}$ is then complemented by an exponentially decreasing term, which smoothly switches the combined correction off when its value becomes too large because of increasing internuclear distances:

$$\begin{aligned} V_{\text{rel+QED}} &= V_{\text{poly}}^{\text{rel+QED}} \cdot h, \\ h &= \frac{1}{1 + e^{2\alpha_c(|V_{\text{poly}}^{\text{rel+QED}}| - V_0)}}, \\ \alpha_c &= \ln\left(\frac{100}{\delta} - 1\right) / b_c, \end{aligned} \quad (16)$$

where $\delta = 0.1$, $b_c = 3 \text{ cm}^{-1}$, $V_0 = 6 \text{ cm}^{-1}$.

2.2.2. Adiabatic correction surface

The present adiabatic correction surface is a slightly modified version of the one computed by us previously [1]. Its polynomial part $V_{\text{poly}}^{\text{ad}}$ has again the analytical

form of Equation (3), with a set of 78 non-zero coefficients V_{nmk}^{ad} and a maximum polynomial power of 12, but in this case transformed coordinates

$$\tilde{R}_{ij} = \left[1 - e^{-\beta(R_{ij}/R_e - 1)}\right] / \beta$$

with $\beta = 1.3$ were used instead of the differences ΔR_{ij} . Parameters V_{nmk}^{ad} were obtained by fitting 5591 adiabatic points computed by Pavanello *et al.* [21], corresponding to energies up to $38,000 \text{ cm}^{-1}$, to $V_{\text{poly}}^{\text{ad}}$. The resulting fit gave an rms value of 0.116 cm^{-1} .

Unlike relativistic and QED corrections, the adiabatic one has an effect on H_3^+ energy states in the near-dissociation region, which is comparable with the corresponding BOPES75K accuracy. To extrapolate the adiabatic correction surface to the region above $38,000 \text{ cm}^{-1}$, we took an adiabatic point, which is close to the dissociation limit, and considered its value $V_0^{\text{ad}} = -114.5 \text{ cm}^{-1}$ as a constant asymptote for H_3^+ molecular configurations with two internuclear distances greater than $R_0^{\text{ad}} = 10 a_0$:

$$\begin{aligned} V_{\text{ad}} &= V_{\text{poly}}^{\text{ad}} h_{\text{low}} + V_0^{\text{ad}} h_{\text{up}}, \\ h_{\text{low}} &= \frac{1}{1 + e^{2\alpha_c^{\text{ad}}(R - R_0^{\text{ad}})}}, \\ h_{\text{up}} &= \frac{1}{1 + e^{-2\alpha_c^{\text{ad}}(R - R_0^{\text{ad}})}}, \\ \alpha_c^{\text{ad}} &= \ln\left(\frac{100}{\delta} - 1\right) / b_c^{\text{ad}}, \end{aligned} \quad (17)$$

where $\delta = 0.1$, $b_c^{\text{ad}} = 1 \text{ cm}^{-1}$, and distance R has the same meaning as earlier in Equation (12).

The Fortran files with BOPES75K and global correction surfaces together with files containing their polynomial constants are presented in the supplementary material.

3. Properties of the new global BO PES

In this work we used a set of 40,537 *ab initio* energies computed by Pavanello *et al.* [21], which span energies up to $75,500 \text{ cm}^{-1}$ and are reproduced by the new potential BOPES75K with a standard deviation value of 4.9 cm^{-1} . 786 points (about 2% from the total set of 41,323 geometries) with energy values from $37,090 \text{ cm}^{-1}$ to $75,380 \text{ cm}^{-1}$ were excluded from our calculations, because their inclusion significantly deteriorates the accuracy. About 20% of them have two large (greater than $7 a_0$) internuclear distances and energies up to $42,000 \text{ cm}^{-1}$, and the others correspond to energies above $53,000 \text{ cm}^{-1}$ and have comparatively small bond lengths. The excluded points are nevertheless represented

Table 3. Standard deviations, in cm^{-1} , of PES energy minus the *ab initio* value for the GLH3P potential [21], the DMBE potential of Viegas *et al.* [17], the PES by Velilla *et al.* [18] and the new BOPES75K potential for different energy ranges.

Energy range/ cm^{-1}	$N(E)$	DMBE	GLH3P	Velilla <i>et al.</i>	BOPES75K
0–35,000	5422	30.56	0.0447	22.34	0.0447
0–37,000	6005	60.42	0.0512	22.37	0.0666
37,000–42,000	1841	151.30	11.18	23.68	3.070
42,000–45,000	934	121.90	27.65	21.38	5.611
45,000–50,000	1503	120.71	34.99*	18.32	4.861
50,000–55,000	1690	111.27	246.34*	17.10	4.981
53,000–75,500	29260	84.12	1.4E+07*	5.872	5.443

Notes: The zero energy for all calculations was taken from the GLH3P potential. $N(E)$ denotes the number of *ab initio* points falling in the given energy range. The rms values indicated by a star* correspond to energy regions that are not covered by the GLH3P surface.

reasonably well by BOPES75K, with a standard deviation of 38.1 cm^{-1} and a maximum deviation of about 130 cm^{-1} .

Table 3 and Figure 2 compare the GLH3P potential, the double many-body expansion (DMBE) potential of Viegas *et al.* [17], whose function form is the basis of the GLH3P PES, the PES by Velilla *et al.* and our new BOPES75K for different energy ranges. It is clear that our new BO PES retains the accuracy of the GLH3P fit at low energy while greatly improving its behaviour near and above dissociation.

Some two-dimensional cuts of our new global *ab initio* BOPES75K for H_3^+ are pictured in Figure 3. They demonstrate the smooth behaviour of our new BO PES, in contrast to the analogous cuts through the GLH3P potential, which show serious unphysical features.

Some two-dimensional cuts and contour plots of our new global *ab initio* BOPES75K for H_3^+ are shown in

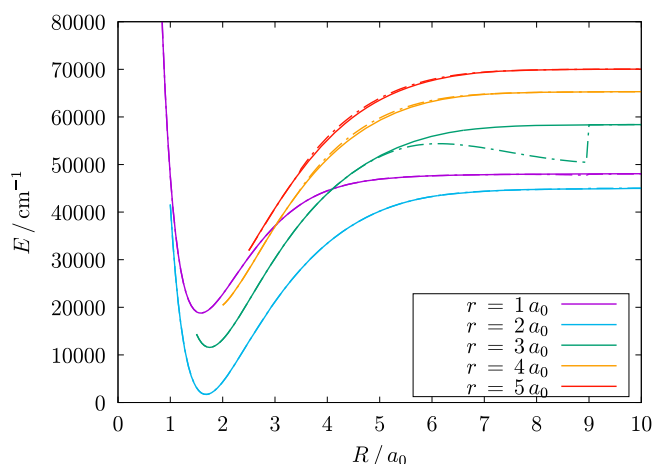


Figure 3. Comparison of some isosceles two-dimensional cuts of the new global *ab initio* BOPES75K for H_3^+ (solid lines) with the ones of the old GLH3P potential (dashed lines). Here the bond lengths of the H_3^+ molecule are $r_1 = r_2 = r_3 = R$.

Figure 3, which compares with the GLH3P potential, and Figure 4, which illustrates the behaviour in switching regions. The plots demonstrate the smooth behaviour of our new BO PES; in contrast some of the analogous cuts through the GLH3P potential show serious unphysical features.

The new BOPES75K can be used to give the dissociation limit into $\text{H}_3^+ \rightarrow \text{H}_2 + \text{H}^+ D_e = 37,195.3 \text{ cm}^{-1}$. This is due to the procedure of obtaining the new BO potential presented in Section 2. For $\text{H}_2 - \text{H}^+$ distances larger than $R = 19 a_0$, the absolute values of differences between the BOPES75K and the original potential by Velilla *et al.* are less than 0.005 cm^{-1} . In particular, the

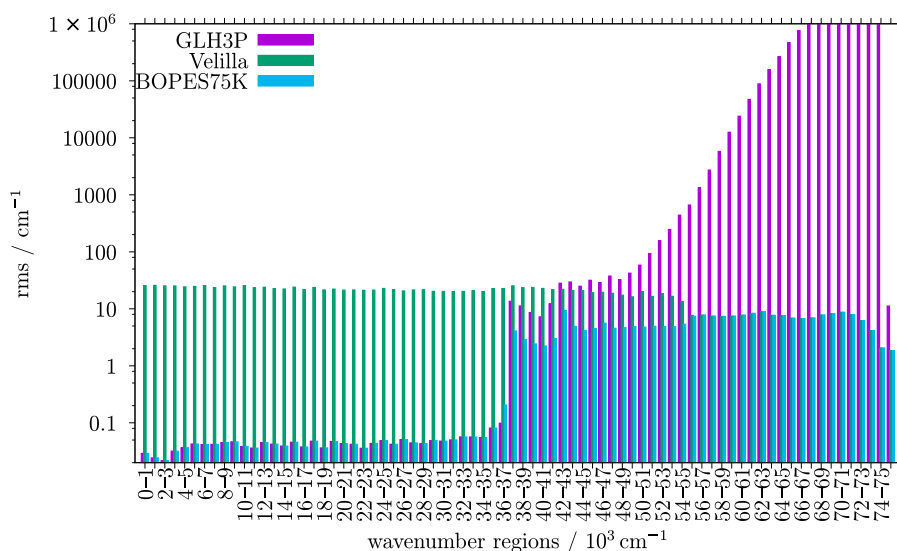


Figure 2. Standard deviations, in cm^{-1} , with which various potentials reproduce the high accuracy *ab initio* points computed by Pavanella *et al.* [21]: GLH3P potential [21], the PES by Velilla *et al.* [18] and the new BOPES75K potential. Energy ranges all have equal widths of 1000 cm^{-1} .

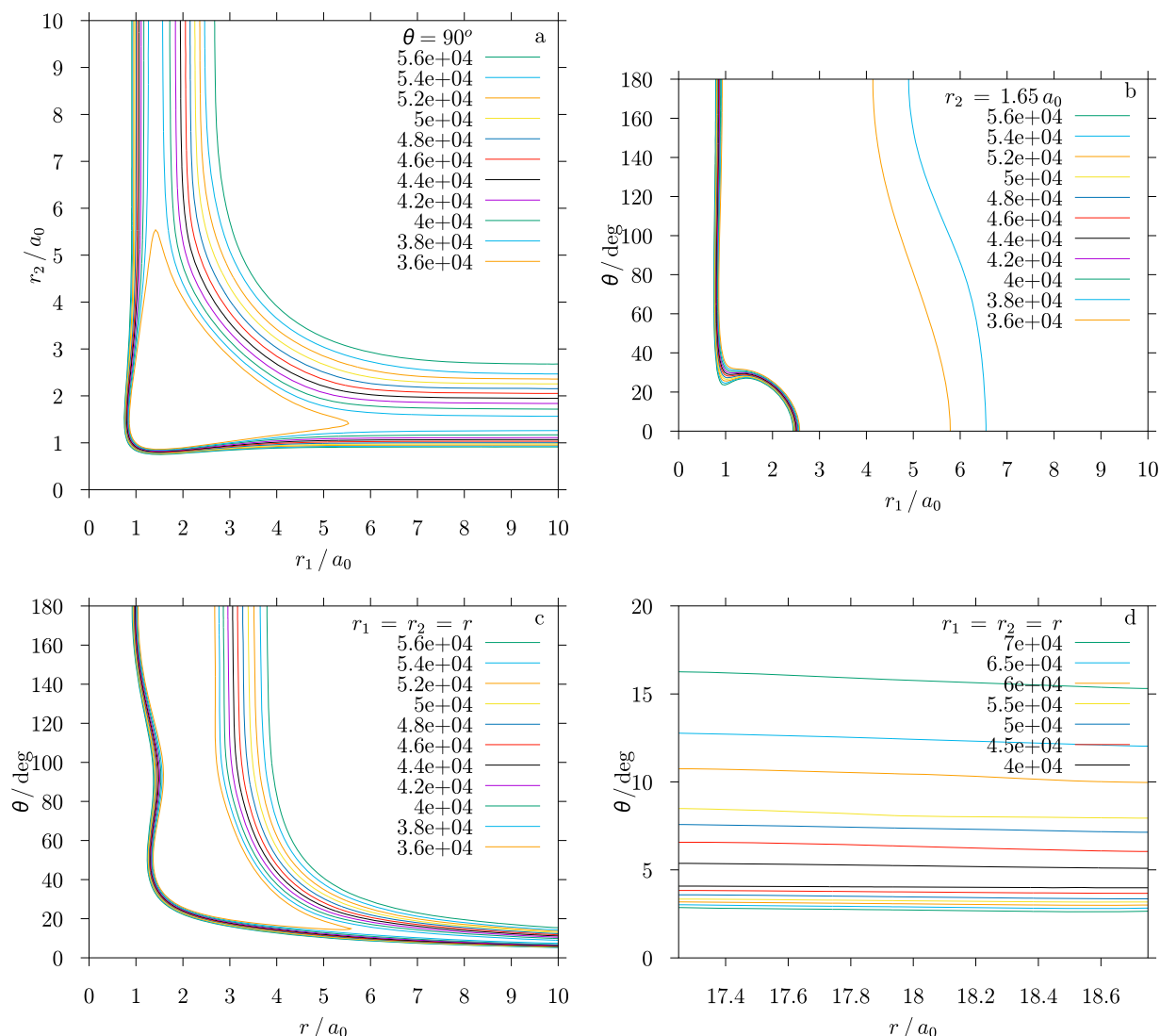


Figure 4. Contour plots of BOPES75K in the switching regions using H – H bondlengths (r_1, r_2) and the angle between them (θ): (a) configurations with $\theta = 90^\circ$ in the energy switching region from 36,000 to 56,000 cm^{-1} ; (b) configurations with $r_2 = 1.65 a_0$, in the same energy switching region; (c) isosceles configurations with $r_1 = r_2 = r$, in the same energy switching region; (d) isosceles configurations with $r_1 = r_2 = r$, for asymptotic switching region with r varied from 17.25 to 18.75 a_0 . Energy values, which correspond to various isolines, are indicated on plots in cm^{-1} .

PES of Velilla *et al.* reproduces a set of high-precision *ab initio* points obtained by them in [18] with a standard deviation of only 1.80 cm^{-1} for $\text{H}_2 - \text{H}^+$ distances R larger than 10 a_0 in the region of the first dissociation limit, and gives a value of the BO dissociation energy $D_e = 37,170 \text{ cm}^{-1}$.

The BO value of dissociation energy D_0 obtained on the basis of our new BOPES75K is $D_0 = 35,011.8 \text{ cm}^{-1}$. In this calculation an experimental value $E_0(\text{H}_2) = 2179.3 \pm 0.1 \text{ cm}^{-1}$ for H_2 vibrational zero-point energy (ZPE) was used [32]. The vibrational ZPE for H_3^+ $E_0(\text{H}_3^+) = 4362.76 \text{ cm}^{-1}$ was obtained from the large basis set calculation performed in Section 4 which is enough to converge the ZPE within 0.0001 cm^{-1} .

For comparison with experiment a further 64.21 cm^{-1} [33] must be added to the ZPE as the Pauli Principle means that the lowest allowed state of H_3^+ has is the $J=1, K=1$. The influence of disregarding non-BO effects is mainly conditioned by the adiabatic correction: the error from neglecting relativistic and QED effects is probably less than 1 cm^{-1} (as in calculations for the H_2 molecule [34]), and the value of nonadiabatic correction is negligible at the dissociation threshold. Our results for the adiabatic correction at equilibrium give a value about 115.1 cm^{-1} , whereas its value near dissociation was taken equal to be 114.5 cm^{-1} (see Section 2.2). Thus, our estimate of the uncertainty in our calculated dissociation energy is about 2 cm^{-1} ,

and $D_0 = 35,076 \pm 2 \text{ cm}^{-1} = 4.3489 \pm 0.0002 \text{ eV}$. This value is a little higher than the best previously-available theoretical result due to Lie and Frye [35] of $4.337 \pm 0.002 \text{ eV}$. However, it remains lower than the best available experiment estimate of $D_0 = 4.381 \pm 0.021 \text{ eV}$ due to Cosby and Helm [36]. We believe our value is the best available estimate of the dissociation energy of H_3^+ .

4. Nuclear motion calculations

In order to test the new PES75K+ surface, calculations were performed using the DVR3D variational nuclear motion programme suite [37]. The calculations were performed in Jacobi coordinates, and the discrete variable representation (DVR) grids were based on spherical oscillator functions for both the atom – diatom coordinate, R , and the diatomic coordinate, r , defined by the parameters $\alpha = 0.0$ and $\omega_e = 0.07$ atomic units [38]. A DVR in (associated) Legendre functions was used for the angular coordinate, θ . The grids contained 60,58, and 68 points for R, r and θ coordinates, respectively. The final diagonalised matrices for the vibrational problem had a dimension of 20,000.

We compute vibrational energy levels, i.e. with total angular momentum $J = 0$, using the new BOPES75K and the GLH3P PES for two cases: (i) for the surfaces in BO approximation using nuclear masses only – up to $60,000 \text{ cm}^{-1}$, and (ii) for these PESs augmented by adiabatic, relativistic and quantum electrodynamics correction surfaces, and allowing for non-adiabatic effects by using different effective vibrational and rotational masses as was suggested by Moss [39]. Calculations were performed up to the first dissociation limit of H_3^+ at about $37,200 \text{ cm}^{-1}$, and a bit above. This non-adiabatic model has been shown to give highly accurate predictions of H_3^+ spectra [40]. In the latter calculation set with BOPES75K the new global correction surfaces with exponential ‘tails’ were used, and the vibrational mass was to 1.007517 Da – a value intermediate between nuclear and atomic masses, which is the optimal one for the issue of the most accurate prediction of experimental vibrational band origins (see Table 6) and was obtained manually by trial-and-error method. For calculations with GLH3P adiabatic and relativistic correction surfaces [21], and a QED correction surface [31], which do not have global character, were used. In these calculations, the value of the vibrational mass derived by Moss for H_2^+ [39] and used previously for H_3^+ [21, 40] was also employed. For calculations with $J > 0$ a rotational mass is also needed; this was always set to the proton (nuclear) mass, as before [21, 40].

Table 4. Root mean square deviations (rms) and maximum absolute deviation (max) values in cm^{-1} of the vibrational (with $J = 0$) Born-Oppenheimer energies obtained using the new global *ab initio* potential BOPES75K from the ones calculated with the GLH3P PES for different energy ranges.

Energy range/ cm^{-1}	$N_v(E)$	rms	max
0–5000	6	0.0000	0.0000
5000–10,000	21	0.0000	0.0001
10,000–15,000	52	0.0001	0.0002
15,000–20,000	110	0.0001	0.0007
20,000–25,000	203	0.0004	0.0015
25,000–30,000	331	0.0017	0.0067
30,000–35,000	528	0.0070	0.0340
35,000–40,000	776	0.0225	0.1924
40,000–45,000	1043	0.2612	1.5575
45,000–50,000	1310	1.8626	8.2676
50,000–55,000	1583	3.8516	20.0573
55,000–60,000	1832	4.4036	14.2542

Note: $N_v(E)$ gives the number of vibrational states given in each energy range.

Table 5. Root mean square deviations (rms) and maximum absolute deviation (max) values in cm^{-1} of vibrational (with $J = 0$) energy levels obtained using the new global *ab initio* potential PES75K+ from the ones calculated with the GLH3P PES, for different energy ranges.

Energy range/ cm^{-1}	$N_v(E)$	rms	max
0–5000	6	0.0770	0.0983
5000–10,000	21	0.1485	0.1885
10,000–15,000	52	0.2059	0.2524
15,000–20,000	110	0.2410	0.2844
20,000–25,000	203	0.2175	0.3291
25,000–30,000	331	0.1281	0.4733
30,000–35,000	528	1.6869	7.0708
35,000–38,000	432	3.5309	10.0955

Notes: Here adiabatic, relativistic, quantum electrodynamics and (partially) nonadiabatic effects were considered in following ways: for PES75K+ – as described above, for GLH3P PES – as in Refs. [21] (adiabatic, relativistic and nonadiabatic parts) and [31] (QED part). $N_v(E)$ gives the number of vibrational states given in each energy range.

The levels of H_3^+ included in the calculations of Table 4 consist of two distinct sets – bound and unbound levels. The bound levels have a clear physical meaning, the unbound ones – the levels above the dissociation, are artefacts of the chosen basis set. However, since the basis sets in both GLH3P and PES75K calculations are the same, the discrepancies between these artefact levels display the real difference between the resonances, which could be obtained using the same PESs and, for example, a complex absorbing potential (CAP) [41, 42].

A comparison between the vibrational energies obtained in the PES75K and GLH3P calculations is performed in Tables 4 and 5. One can see that there is a minor difference between the energies calculated on the basis of the two BO potentials up to the dissociation energy value, but significant differences (up to tens of cm^{-1}) appear in energy range from about $40,000 \text{ cm}^{-1}$ and above. This is a direct consequence of fixing the unsmooth parts of the GLH3P PES.

Table 6. Comparison of vibrational band origins (energy levels with $J = 0$) obtained using the new global *ab initio* potential PES75K+ with the available experimental data [43].

Symmetry	$E_{\text{exp}}/\text{cm}^{-1}$	PES75K+/ cm^{-1}	Obs. -					
			calc./ cm^{-1}	K	G	ν_1	ν_2	l_2
E'	2521.411	2521.307	0.104	0	1	0	1	1
E'	4998.049	4997.912	0.137	0	2	0	2	2
E'	5554.060	5554.223	-0.163	0	1	1	1	1
A_2'	7492.912	7492.807	0.106	0	3	0	3	3
E'	9113.080	9113.077	0.003	0	2	0	4	2
E'	10645.380	10645.338	0.042	0	2	2	2	2
E'	10862.910	10862.780	0.130	0	1	0	5	1
E'	11323.100	11323.145	-0.045	0	1	3	1	1
E'	11658.400	11658.344	0.056	0	5	0	5	5
E'	12303.370	12303.376	-0.006	0	1	2	3	1
E'	12477.380	12477.432	-0.052	0	2	0	6	2
E'	13702.380	13702.676	-0.296	0	1	0	7	1
E'	15122.810	15122.725	0.085	0	2	0	8	2

Notes: Adiabatic, relativistic, quantum electrodynamics and (partially) non-adiabatic effects were considered in this set of calculations. The calculated levels with symmetry E' were obtained as an arithmetic mean of (quasi-) degenerate even and odd pairs of DVR3D levels, which correspond to degenerate f-symmetry states. The experimental energies are supplemented by corresponding quantum numbers sets. Here (ν_1, ν_2) are vibrational quantum numbers, K is an absolute value of the projection of J on the C_3 axis and G is an absolute value of quantum number $g = k - l_2$.

Thus, Table 5 reflects only changes in the way non-BO effects are taken into account as it covers only the energy region up to $37,000 \text{ cm}^{-1}$. The difference between the approaches performed in the present work and previously in works by Pavanello *et al.* [21] and Lodi *et al.* [31] becomes significant for energies above $30,000 \text{ cm}^{-1}$, i.e. in the energy region where the old correction surfaces were not fitted accurately.

Finally, in Table 6 a comparison of vibrational energy levels obtained using PES75K+ with adiabatic, relativistic, quantum electrodynamics, and (partially) nonadiabatic corrections with the available experimental data for states with $J=0$ is performed. The standard deviation obtained in this comparison is about 0.12 cm^{-1} , which is almost half the value 0.21 cm^{-1} obtained with vibrational masses suggested by Moss [39].

5. Conclusions

We present a modified global H_3^+ PES in both a simple BO form and as an augmented BO plus relativistic, QED and adiabatic corrections PES. The *ab initio* points used for this representation [24] are extremely accurate (accurate to an absolute energy of about 0.01 cm^{-1}) and the only problem with the previous fit to these points was that analytical representation of some configurations with energy values above $37,000 \text{ cm}^{-1}$ was poor leading them to be excluded from the fit. The task of representing all the *ab initio* points with the intrinsic accuracy of *ab initio* calculations proved to be difficult, and the improved PESs presented here still do not recover the full

accuracy of the underlying *ab initio* electronic structure points. Our new representation corrects the shortcomings of the previous incarnation [21] and displays the correct asymptotic behaviour.

Our new PES is suitable for a variety of calculations. Firstly, scattering calculations for collisions between protons and diatomic hydrogen molecules, particularly at ultra-low energies where such collisions are very sensitive to the details of the underlying PES. Secondly, the study of asymptotic vibrational states for the H_3^+ system [44]. Thirdly, for accurate calculation of H_3^+ resonance states, which should finally lead to the interpretation of the famous, indeed notorious, Carrington-Kennedy predissociation spectrum [45, 46].

Acknowledgments

We thank Eryn Spinlove and Attila Császár for comments on our potential energy surface.

Disclosure statement

No potential conflict of interest was reported by the authors.

Funding

This work was partially supported by the Russian Fund for Basic Research (RFBR) as part of the research project # 18-02-00705 and by the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant through grant number 701962.

ORCID

Irina I. Mizus  <http://orcid.org/0000-0001-6394-0129>

Jonathan Tennyson  <http://orcid.org/0000-0002-4994-5238>

Alexander Alijah  <http://orcid.org/0000-0002-4915-0558>

References

- [1] J. Tennyson, O.L. Polyansky, N.F. Zobov, A. Alijah and A.G. Császár, *J. Phys. B: At. Mol. Opt. Phys.* **50**, 232001 (2017).
- [2] T. Gonzalez-Lezana and P. Honvault, *Intern. Rev. Phys. Chem.* **33**, 371 (2014).
- [3] A. Carrington, J. Buttenshaw and R. Kennedy, *Mol. Phys.* **45**, 753 (1982).
- [4] A. Carrington and R.A. Kennedy, *J. Chem. Phys.* **81**, 91 (1984).
- [5] A. Carrington, *J. Chem. Soc. Farad. T.2* **82**, 1089 (1986).
- [6] A. Carrington and I.R. McNab, *Acc. Chem. Res.* **22**, 218 (1989).
- [7] A. Carrington, I.R. McNab and Y.D. West, *J. Chem. Phys.* **98**, 1073 (1993).
- [8] F. Kemp, C.E. Kirk and I.R. McNab, *Phil. Trans. A* **358** (1774), 2403 (2000).
- [9] N.F. Zobov, O.L. Polyansky, C.R. Le Sueur and J. Tennyson, *Chem. Phys. Lett.* **260**, 381 (1996).

- [10] E. Carmona-Novillo, T. Gonzalez-Lezana, O. Roncero, P. Honvault, J.M. Launay, N. Bulut, F.J. Aoiz, L. Banares, A. Trotter and E. Wrede, *J. Chem. Phys.* **128**, 014304 (2008).
- [11] P. Honvault and Y. Scribano, *J. Phys. Chem. A* **117**, 9778 (1997).
- [12] M. Lara, P.G. Jambrina, F.J. Aoiz and J.M. Launay, *J. Chem. Phys.* **143**, 204305 (2015).
- [13] A. Ichihara and K. Yokoyama, *J. Chem. Phys.* **103**, 2109 (1995).
- [14] R. Prosimiti, O.L. Polyansky and J. Tennyson, *Chem. Phys. Lett.* **273**, 107 (1997).
- [15] O.L. Polyansky, R. Prosimiti, W. Klopper and J. Tennyson, *Mol. Phys.* **98**, 261 (2000).
- [16] A. Aguado, O. Roncero, C. Tablero, C. Sanz and M. Paniagua, *J. Chem. Phys.* **112**, 1240 (2000).
- [17] L.P. Viegas, A. Alijah and A.J.C. Varandas, *J. Chem. Phys.* **126**, 074309 (2007).
- [18] L. Velilla, B. Lepetit, A. Aguado, J.A. Beswick and M. Paniagua, *J. Chem. Phys.* **129**, 084307 (2008).
- [19] R.A. Bachorz, W. Cencek, R. Jaquet and J. Komasa, *J. Chem. Phys.* **131**, 024105 (2009).
- [20] P. Barragan, R. Prosimiti, P. Villarreal and G. Delgado-Barrio, *Int. J. Quant. Chem.* **111**, 368 (2011).
- [21] M. Pavanello, L. Adamowicz, A. Alijah, N.F. Zobov, I.I. Mizus, O.L. Polyansky, J. Tennyson, T. Szidarovszky and A.G. Császár, *J. Chem. Phys.* **136**, 184303 (2012).
- [22] R. Jaquet and M.V. Khoma, *J. Chem. Phys.* **136**, 154307 (2012).
- [23] B. Mukherjee, S. Mukherjee and S. Adhikari, *J. Phys. Conf. Ser.* **759** (1), 012050 (2016).
- [24] M. Pavanello, L. Adamowicz, A. Alijah, N.F. Zobov, I.I. Mizus, O.L. Polyansky, J. Tennyson, T. Szidarovszky, A.G. Császár, M. Berg, A. Petignani and A. Wolf, *Phys. Rev. Lett.* **108**, 023002 (2012).
- [25] O.L. Polyansky, A. Alijah, N.F. Zobov, I.I. Mizus, R. Ovsyannikov, J. Tennyson, T. Szidarovszky and A.G. Császár, *Phil. Trans. Royal Soc. London A* **370**, 5014 (2012).
- [26] M. Pavanello and L. Adamowicz, *J. Chem. Phys.* **130**, 034104 (2009).
- [27] A.J.C. Varandas, *J. Chem. Phys.* **105**, 3524 (1996).
- [28] A.J.C. Varandas, *J. Chem. Phys.* **119**, 2596 (2003).
- [29] W.R. Gentry and C.F. Giese, *Phys. Rev. A* **11**, 90 (1975).
- [30] R.A. Bachorz, W. Cencek, R. Jaquet and J. Komasa, *J. Chem. Phys.* **131**, 024105 (2009).
- [31] L. Lodi, O.L. Polyansky, J. Tennyson, A. Alijah, and N.F. Zobov, *Phys. Rev. A* **89**, 032505 (2014).
- [32] K.K. Irikura, *J. Phys. Chem. Ref. Data* **36**, 389 (2007).
- [33] J. Ramanlal, O.L. Polyansky and J. Tennyson, *Astron. Astrophys.* **406**, 383 (2003).
- [34] M. Puchalski, J. Komasa, P. Czachorowski and K. Pachucki, *Phys. Rev. Lett.* **117**, 263002 (2016).
- [35] G.C. Lie and D. Frye, *J. Chem. Phys.* **96**, 6784 (1992).
- [36] P.C. Cosby and H. Helm, *Chem. Phys. Lett.* **152**, 71 (1988).
- [37] J. Tennyson, M.A. Kostin, P. Barletta, G.J. Harris, O.L. Polyansky, J. Ramanlal and N.F. Zobov, *Comput. Phys. Commun.* **163**, 85 (2004).
- [38] J. Tennyson and B.T. Sutcliffe, *J. Mol. Spectrosc.* **101**, 71 (1983).
- [39] R.E. Moss, *Mol. Phys.* **89**, 195 (1996).
- [40] O.L. Polyansky and J. Tennyson, *J. Chem. Phys.* **110**, 5056 (1999).
- [41] U.V. Riss and H.D. Meyer, *J. Phys. B: At. Mol. Opt. Phys.* **26**, 4503 (1993).
- [42] B.C. Silva, P. Barletta, J.J. Munro and J. Tennyson, *J. Chem. Phys.* **128**, 244312 (2008).
- [43] T. Furtenbacher, T. Szidarovszky, E. Mátyus, C. Fábri and A.G. Császár, *J. Chem. Theory Comput.* **9**, 5471 (2013).
- [44] J.J. Munro, J. Ramanlal and J. Tennyson, *New J. Phys.* **7**, 196 (2005).
- [45] A. Carrington, J. Buttenshaw and R.A. Kennedy, *Mol. Phys.* **45**, 753 (1982).
- [46] A. Carrington and R.A. Kennedy, *J. Chem. Phys.* **81**, 91 (1984).